Department of Physics Indian Institute of technology Madras Select/Special Topics in Classical Mechanics Self-Assessment-2 (Questions & Answers)

<u>NOTE</u>: Symbols/notations used in this question paper have their usual meanings, as used in our course.

\odot SOLUTIONS \odot

- 1. State whether the following statements are 'TRUE' or 'FALSE' <u>and give reason</u>. The reason should be short, but as rigorous as you can provide.
 - **a.** For a particle of mass m moves in a region of space where the potential is described

by $U(x,y) = -U_0 \exp\left[-\frac{(x^2 + y^2)}{2L^2}\right]$, the point (x=0, y=0) is a 'saddle point' (given: $U_0 \& L$ are

positive constants). Solution: False

$$\frac{d^2U}{dx^2} = \frac{U_0}{L^2} \quad \text{at point } (x, y) = (0, 0)$$

$$\frac{d^2U}{dy^2} = \frac{U_0}{L^2} \quad \text{at point}(x, y) = (0, 0)$$

Since double derivative of the function with respect to the two variable is positive, the point at (x,y) = (0,0) is not a saddle point; it is a point of 'stable equilibrium'.

b. If a vector field \vec{A} is both irrotational $(\vec{\nabla} \times \vec{A} = \vec{0})$ and solenoidal $(\vec{\nabla} \cdot \vec{A} = 0)$, then it must be identically equal to the *null vector*.

Solution: False For any constant vector field \vec{A} , for example, $(\vec{\nabla} \times \vec{A} = \vec{0})$ and $(\vec{\nabla} \cdot \vec{A} = 0)$ 2. A position-dependent force field is given by the expression $\vec{F} = A(x-y)\hat{e}_x + (x+y)\hat{e}_y$. It is given that |A| = +1.

(a) What is/are the dimension(s) of A?

Solution:

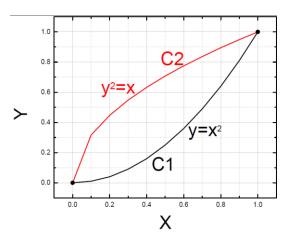
Dimension of A must be
$$\frac{\lfloor F \rfloor}{L} = MT^{-2}$$

(b) The given force acts on a particle, moving it along a closed path described by the two curves:

 $y = x^2$, traversed from (0,0) to (1,1),

and

 $y^{2} = x$ traversed from (1, 1) to (0, 0).



(c) Determine the work $\oint \vec{F} \cdot \vec{dI}$ done by the above force over the closed path described above.

Solution:

$$\vec{F} = A(x-y)\hat{e}_{x} + (x+y)\hat{e}_{y} \text{ with } |A|=1$$

$$\oint_{C} \vec{F} \cdot d\vec{l} = \int_{Along C_{1}} \vec{F} \cdot d\vec{l} + \int_{Along C_{2}} \vec{F} \cdot d\vec{l}$$

$$\int_{Along C_{1}} \vec{F} \cdot d\vec{l} = \int_{0}^{1} (x-x^{2})dx + (x+x^{2})2xdx = \frac{4}{3}$$

$$\int_{Along C_{2}} \vec{F} \cdot d\vec{l} = \int_{1}^{01} (y^{2}-y)2ydy + (y^{2}+y)dy = -\frac{2}{3}$$

$$\therefore \oint_{C} \vec{F} \cdot d\vec{l} = \int_{Along C_{1}} \vec{F} \cdot d\vec{l} + \int_{Along C_{2}} \vec{F} \cdot d\vec{l} = \frac{2}{3}$$

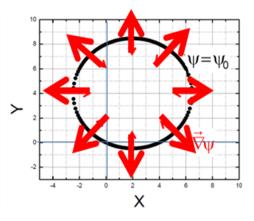
(d) <u>Without</u> determining the curl of this force (i.e. without finding $\vec{\nabla} \times \vec{F}$, can you tell if the force is irrotational or not? Explain how!

Solution:

The Stokes' theorem states that $\oint \vec{F} \cdot \vec{dl} = \iint (\vec{\nabla} \times \vec{F}) \cdot \vec{ds}$. In the present case, since $\oint \vec{F} \cdot \vec{dl}$ is nonzero,

 $\vec{\nabla} \times \vec{F}$ must also be nonzero, which implies that \vec{F} is not irrotational.

- 3. A scalar field $\psi(x, y)$ is given by the expression $\psi(x, y) = \psi_0 \exp(x^2 + y^2 4x 8y)$, where ψ_0 is a constant having suitable dimensions.
 - (a) Obtain the equipotential curve for $\psi = \psi_0$. Solution:
 - $\psi(x, y) = \psi_0 \exp(x^2 + y^2 4x 8y)$ $\ln \frac{\psi}{\psi_0} = x^2 + y^2 - 4x - 8y$ Add 20 to both sides, $\ln \frac{\psi}{\psi_0} + 20 = x^2 + y^2 - 4x - 8y + 20$ $= (x - 2)^2 + (y - 4)^2$



For constant value of $\psi = \psi_0$, the equipotential curve is a circle with centre at (x,y)=(2,4) and radius $\sqrt{20} \approx 4.47...$.

b) Sketch the vector field $\vec{\nabla}\psi$ at $\psi = \psi_0$.

Solution: $\nabla \psi$ is perpendicular to equipotential curves, as shown. Note that as $\psi \ge \psi_0$, the radius of the equipotential circle would be $\ge \sqrt{20}$, so the gradient would be pointed OUTWARD.

4 (a) Determine the divergence of the vector point function described by:

 $\vec{A}(\hat{r}) = (r\cos\theta) \hat{e}_r + (r\sin\theta) \hat{e}_{\theta} + (r\sin\theta\cos\phi) \hat{e}_{\phi}$

Solution:

$$\vec{\nabla} \bullet \vec{A} = \left\{ \hat{e}_r \frac{\partial}{\partial r} + \hat{e}_{\theta} \frac{1}{r} \frac{\partial}{\partial \theta} + \hat{e}_{\phi} \frac{1}{r \sin \theta} \frac{\partial}{\partial \phi} \right\} \bullet \left\{ (\mathbf{r} \cos \theta) \, \hat{e}_r + (\mathbf{r} \sin \theta) \, \hat{e}_{\theta} + (\mathbf{r} \sin \theta \cos \theta) \, \phi \, \hat{e}_{\phi} \right\}$$
$$= \frac{1}{r^2} \frac{\partial}{\partial r} (r^2 r \cos \theta) + \frac{1}{r \sin \theta} \frac{\partial}{\partial \theta} (\sin \theta r \sin \theta) + \frac{1}{r \sin \theta} \frac{\partial}{\partial \phi} (r \sin \theta \cos \phi)$$
$$= 3 \cos \theta + 2 \cos \theta + (-\sin \phi) = 5 \cos \theta - \sin \phi$$

(b) Find the flux of the above vector field over a *closed* surface that encloses a hemisphere of radius R resting on the xy-plane, with its center at origin and located in the region $z \ge 0$.

Solution: Net flux=
$$\oint \vec{A} \cdot \vec{ds} = \iint_{\substack{upper\\hemisphere}} \vec{A} \cdot \vec{ds} + \iint_{\substack{creader\\argenter}} \vec{A} \cdot \vec{ds}$$

 $\vec{A} \cdot \vec{ds} = \iint_{argenter} \vec{A} \cdot \vec{ds}$
 $\vec{A} \cdot \vec{e}_z = (r \cos\theta) \hat{e}_r + (r \sin\theta) \hat{e}_{\theta} + (r \sin\theta \cos\phi) \hat{e}_{\theta}$
 $\vec{A} \cdot \hat{e}_z = (r \cos\theta) (\hat{e}_r \cdot \hat{e}_z) + (r \sin\theta) (\hat{e}_{\theta} \cdot \hat{e}_z) +$
 $(r \sin\theta \cos\phi) (\hat{e}_{\theta} \cdot \hat{e}_z)$
 $\hat{e}_r = \sin\theta \cos\phi \hat{e}_x + \sin\theta \sin\phi \hat{e}_y + \cos\theta \hat{e}_z$
 $\hat{e}_{\theta} = \cos\theta \cos\phi \hat{e}_x + \cos\theta \sin\phi \hat{e}_y - \sin\theta \hat{e}_z$
 $\hat{e}_{\phi} = -\sin\phi \hat{e}_x + \cos\phi \hat{e}_y$
 $\vec{A} \cdot \hat{e}_z = (r \cos\theta) (\cos\theta) + (r \sin\theta) (-\sin\theta)$
 $= r(\cos^2\theta - \sin^2\theta) = -r$ at $\theta = \frac{\pi}{2}$ (xy plane)
 $\vec{A} \cdot \vec{a} \cdot$

5 A planet in a remote galaxy rotates rapidly about its own axis. It completes one full rotation in one second. Sketch $T(\lambda)$ vs λ for this planet, where $T(\lambda)$ is the time period for the rotation of a Foucault

