



Solar Energy: The Sun to Earth Transaction

Learning objectives:

- 1) To calculate the energy received by the Earth from the Sun
- 2) To compare the energy received by Earth from the Sun, with the energy usage by humankind

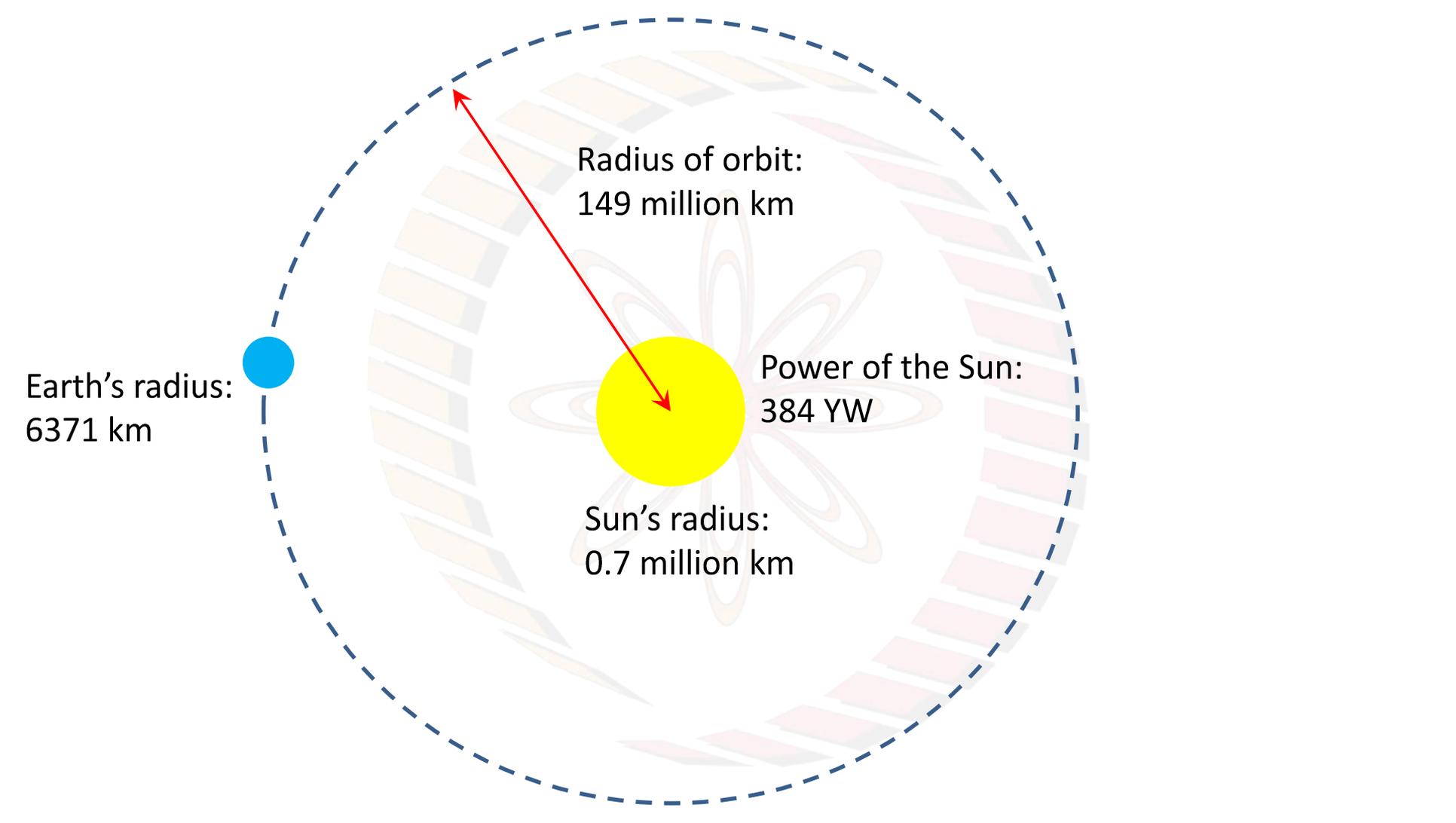
Surface of the Sun ~ 5500 °C

Core of the Sun, several million °C

Sun gives out **384 Yotta Watts**

$$= 384 \times 10^{24} \text{ W}$$

$$= 3.84 \times 10^{26} \text{ W}$$



The diagram illustrates the Sun at the center, represented by a yellow circle. A dashed blue circle represents Earth's orbit. A red arrow points from the Sun to the orbit, indicating the orbital radius. A blue circle on the orbit represents Earth, with a line connecting it to the Sun's center. The background features a faint, stylized sunburst pattern.

Earth's radius:
6371 km

Radius of orbit:
149 million km

Power of the Sun:
384 YW

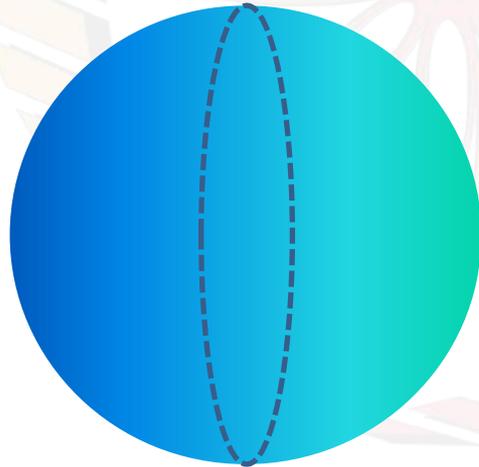
Sun's radius:
0.7 million km

Intensity of Sun's radiation at Earth's orbit:

$$= \frac{3.84 \times 10^{26}}{4 \times 3.14 \times (1.49 \times 10^{11})^2} = 1377 \text{ W/m}^2$$

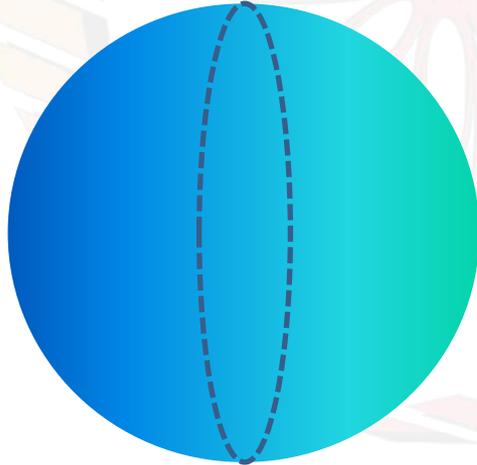
Area of Earth's disc:

$$= 3.14 \times (6.371 \times 10^6)^2 = 1.27 \times 10^{14} \text{ m}^2$$



Power received from the Sun, by Earth:

$$= 1.27 \times 10^{14} \times 1377 = 1.755 \times 10^{17} \text{ W or J/s}$$

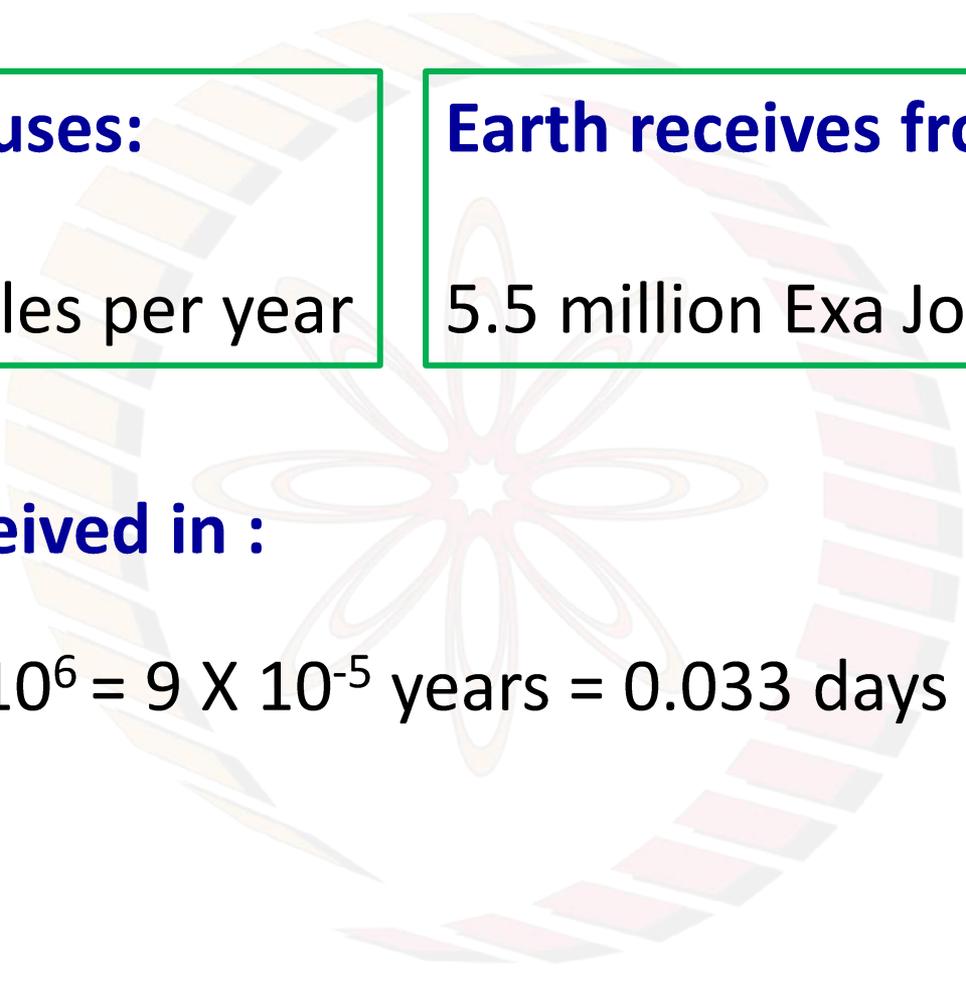


Energy received from the Sun, by Earth each year:

$$= 1.755 \times 10^{17} \times 60 \times 60 \times 24 \times 365$$

$$= 5.5 \times 10^{24} \text{ J}$$

= 5.5 million Exa Joules per year



Humankind uses:

= 500 Exa Joules per year

Earth receives from the Sun:

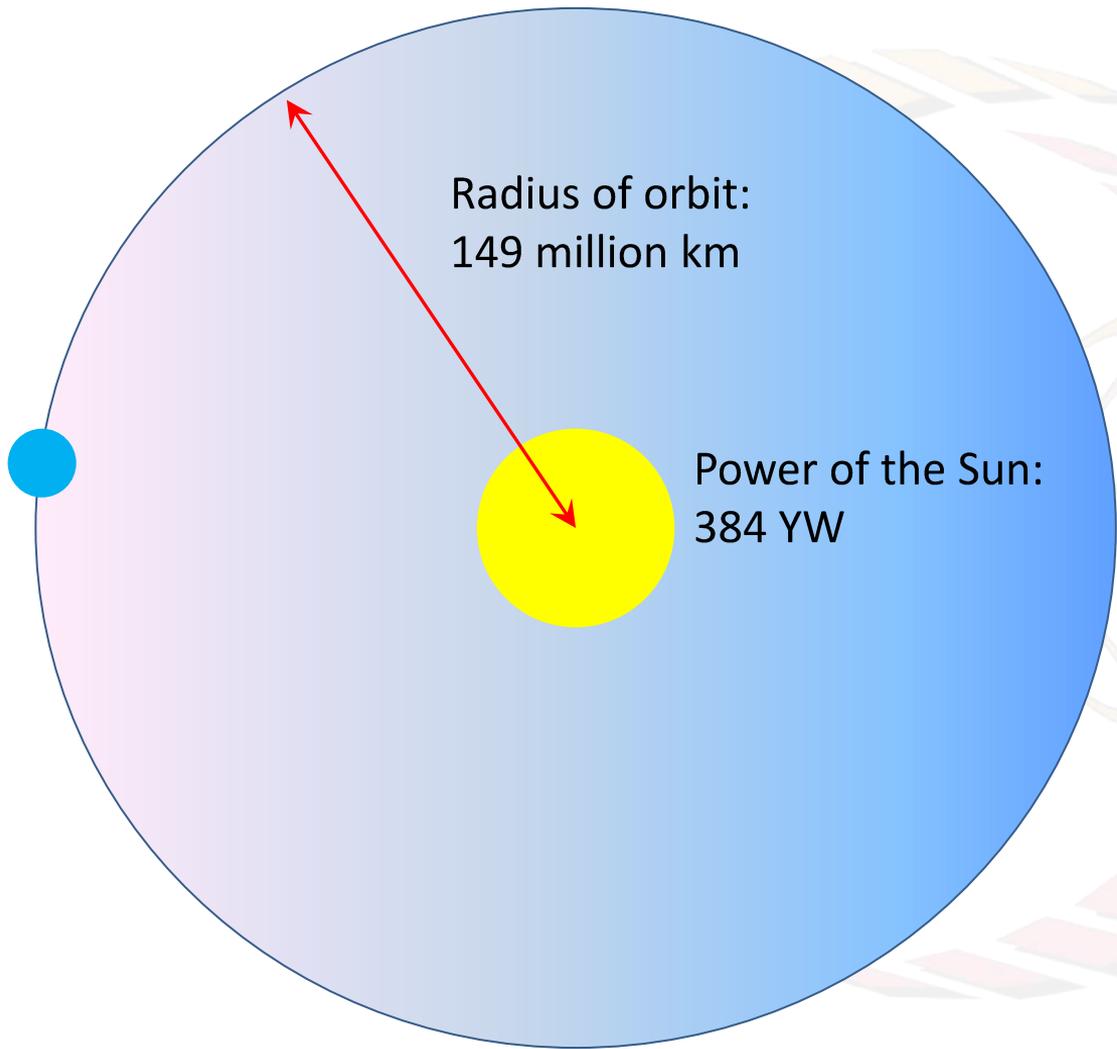
5.5 million Exa Joules per year

This is received in :

$500 / 5.5 \times 10^6 = 9 \times 10^{-5}$ years = 0.033 days = 0.79 hours

Since 30% of the incident energy is reflected back, on the surface of the Earth, the energy used by humankind each year is received in :

$$= 0.79/0.70 \sim \mathbf{1 \text{ hour}}$$



Dyson Sphere:

Freeman Dyson (1960)

Kardashev Scale (1964)

Type 1: All energy reaching planet

Type 2: All energy of Star

Type 3: All energy of Galaxy

Conclusions:

- 1) Earth receives nearly 5.5 million exajoules of energy from the Sun each year
- 2) The entire energy used by humankind each year, is received on the surface of the earth, from the sun, each hour!



The Solar Energy Budget

Learning objectives:

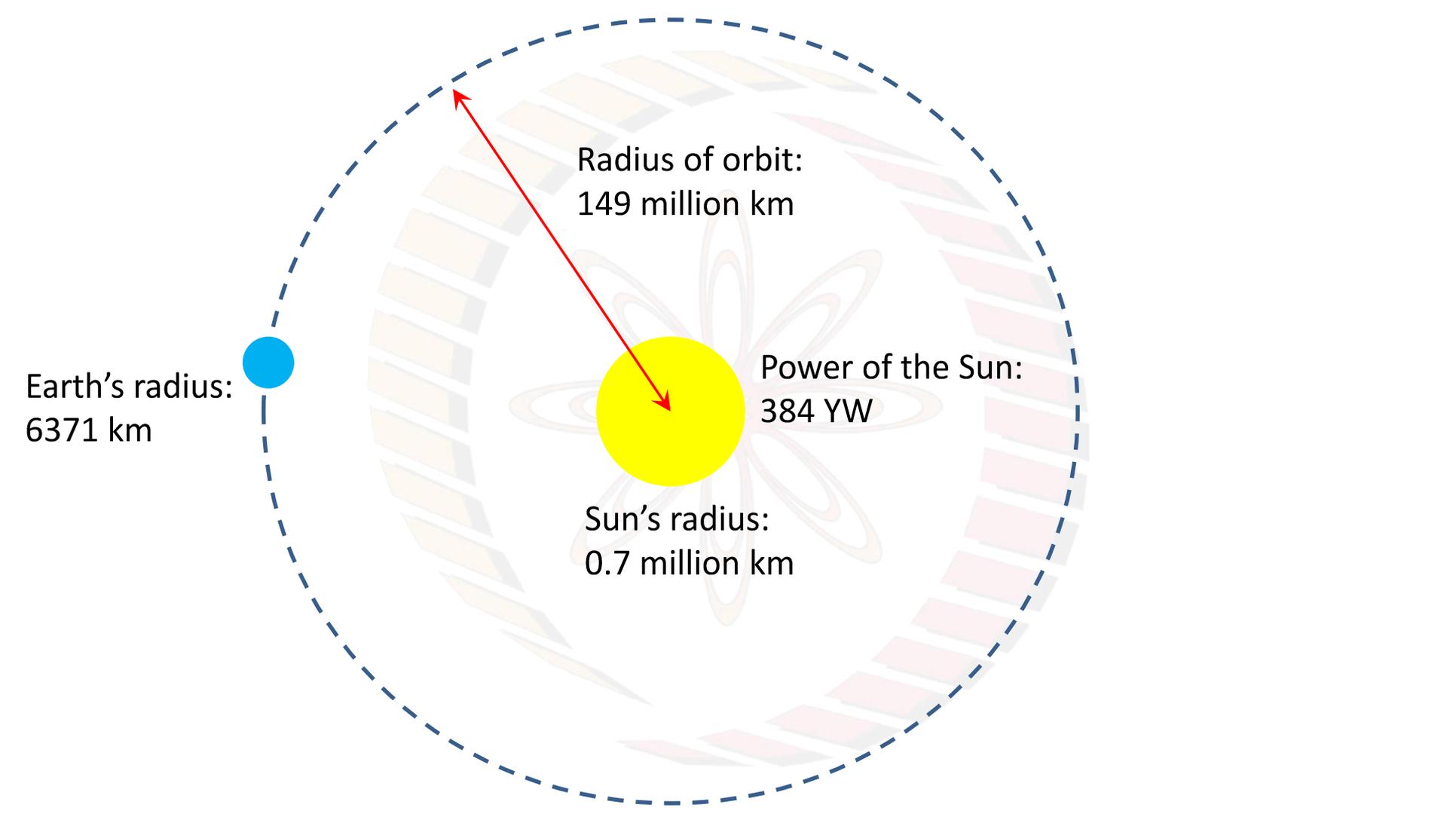
- 1) To know how incoming solar energy is distributed across various phenomena on earth
- 2) To become aware of geographical and seasonal aspects associated with solar energy
- 3) To become aware of impact of time of day on ability to receive solar energy

Layers of the atmosphere:

Exosphere:	700 to 10,000 km
Thermosphere:	80 to 700 km
Mesosphere:	50 to 80 km
Stratosphere:	12 to 50 km
Troposphere:	0 to 12 km

Approximately 80% of mass of Earth's atmosphere is in the Troposphere

Source: https://en.wikipedia.org/wiki/Atmosphere_of_Earth



The diagram illustrates the Sun at the center, represented by a yellow circle. A dashed blue circle represents Earth's orbit. A red arrow points from the Sun to the orbit, labeled 'Radius of orbit: 149 million km'. A blue circle on the orbit represents Earth, with a line pointing to the text 'Earth's radius: 6371 km'. The Sun's radius is labeled as '0.7 million km' and the power of the Sun as '384 YW'. The background features a faint, stylized sunburst pattern.

Earth's radius:
6371 km

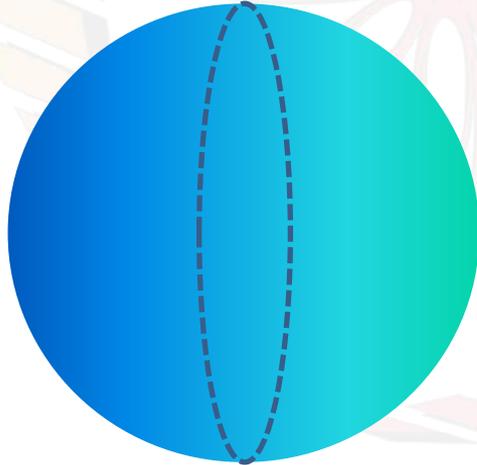
Radius of orbit:
149 million km

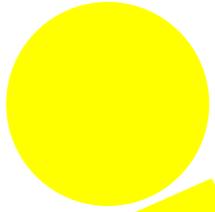
Power of the Sun:
384 YW

Sun's radius:
0.7 million km

Power received from the Sun, by Earth:

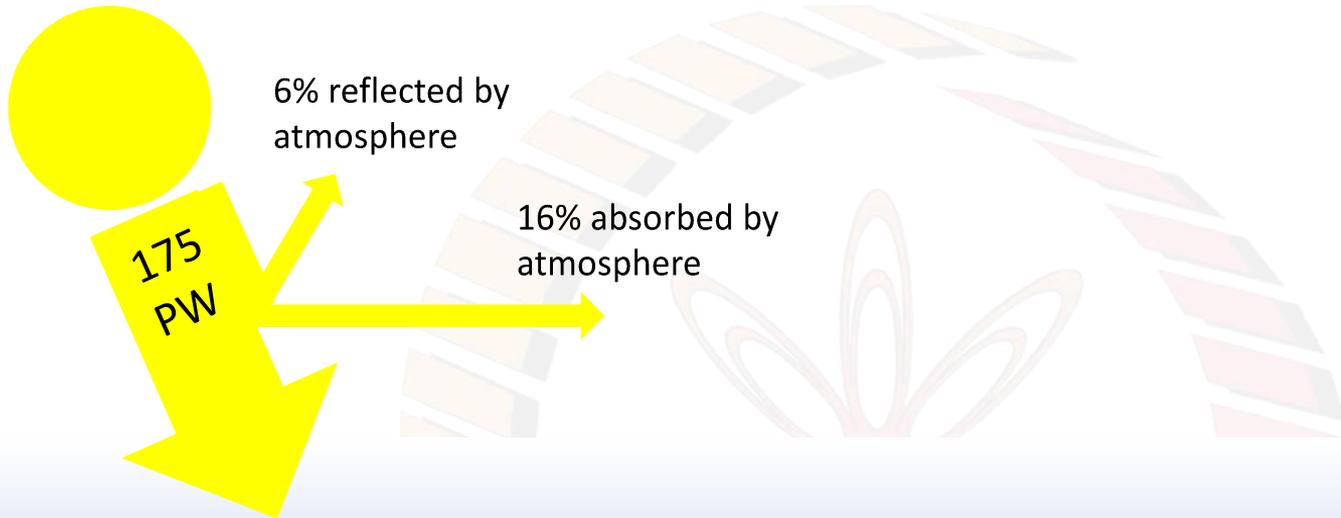
$$= 1.27 \times 10^{14} \times 1377 = 1.755 \times 10^{17} \text{ W or J/s}$$

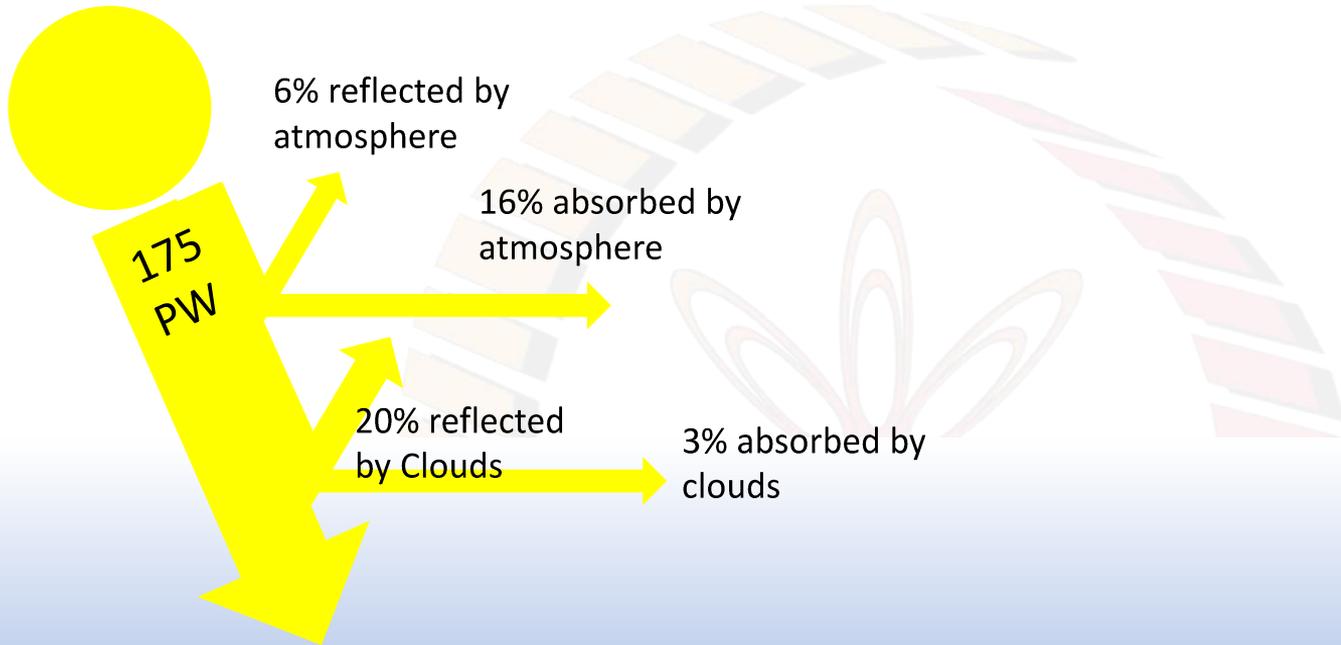


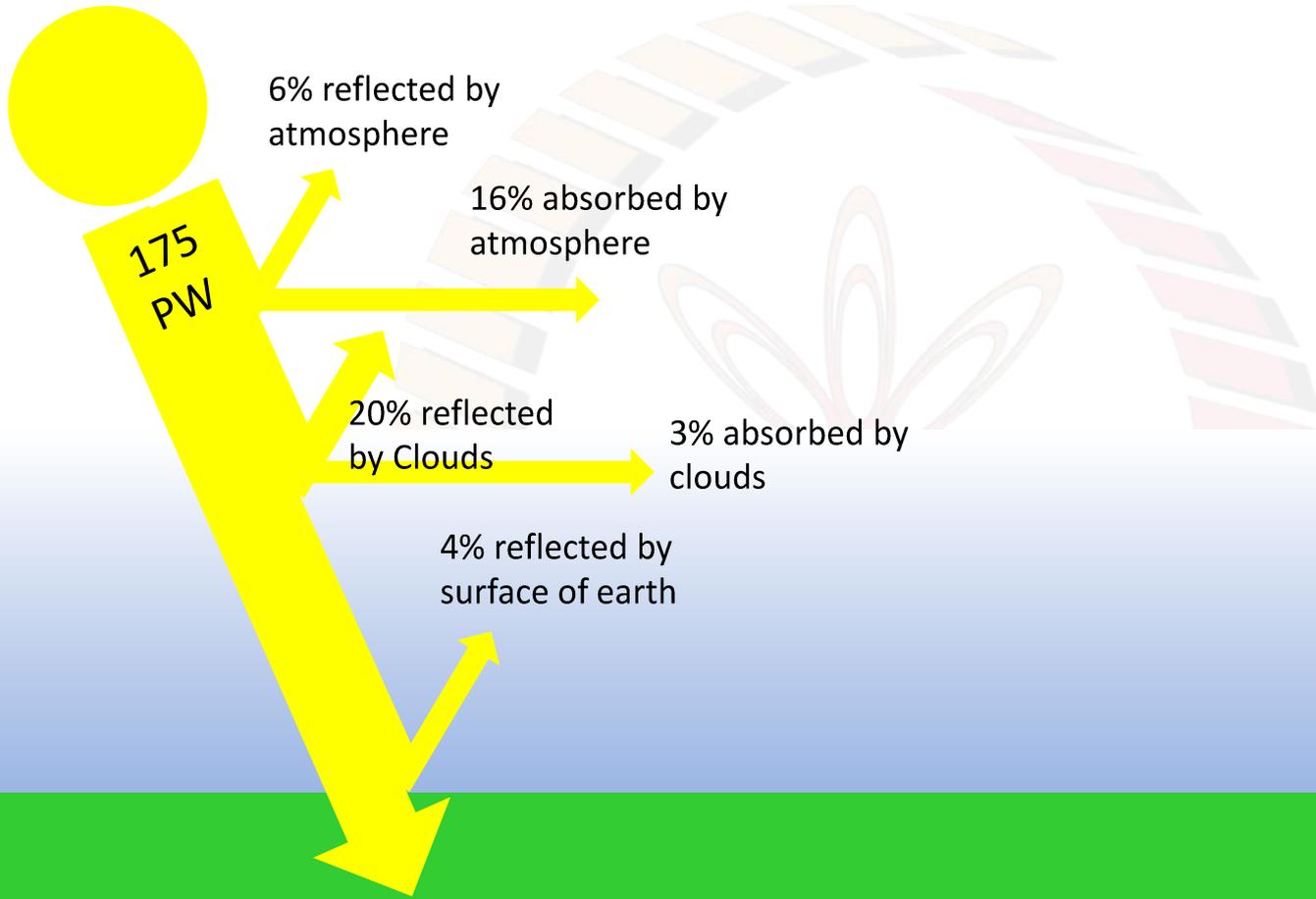


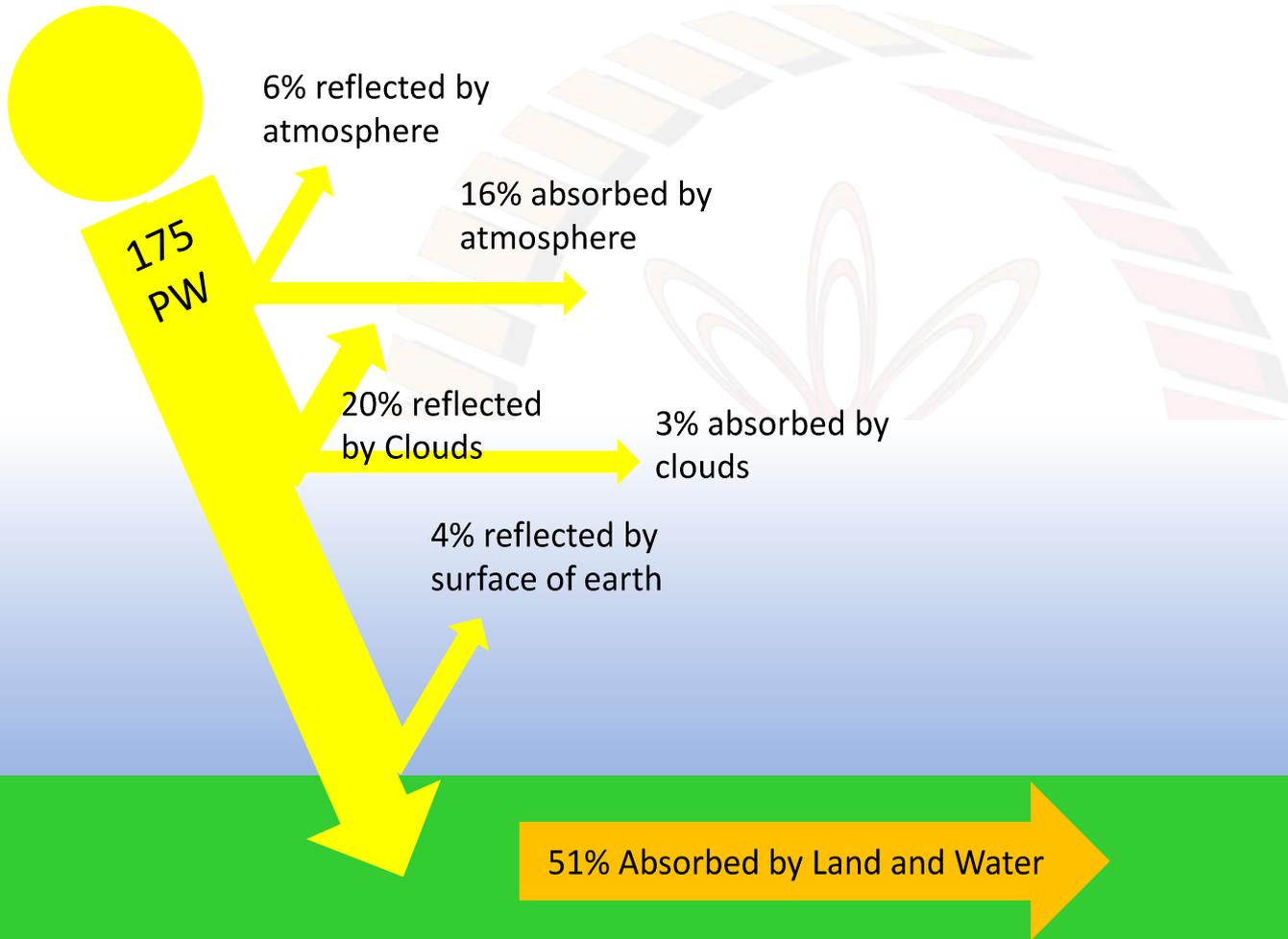
175
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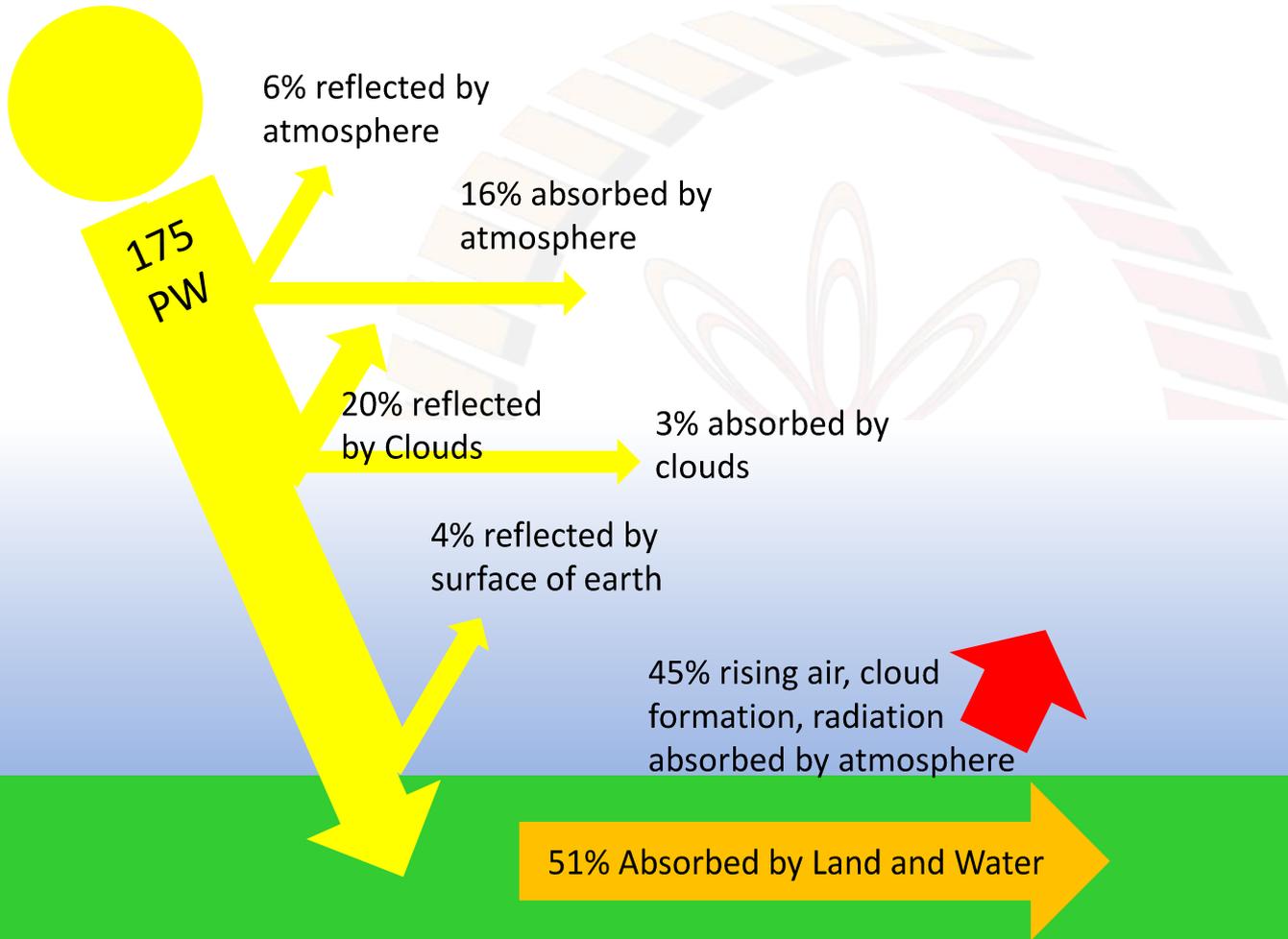


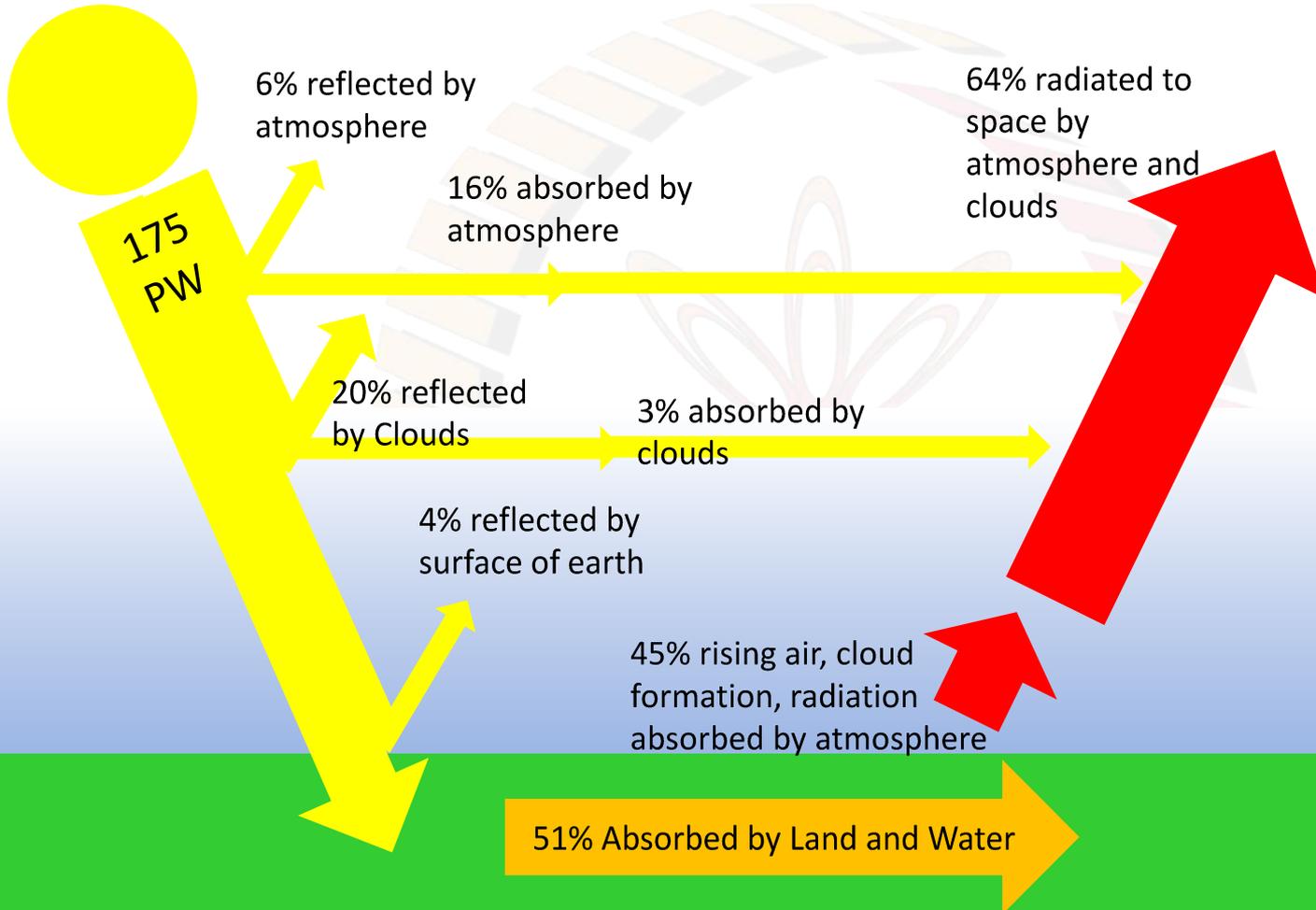


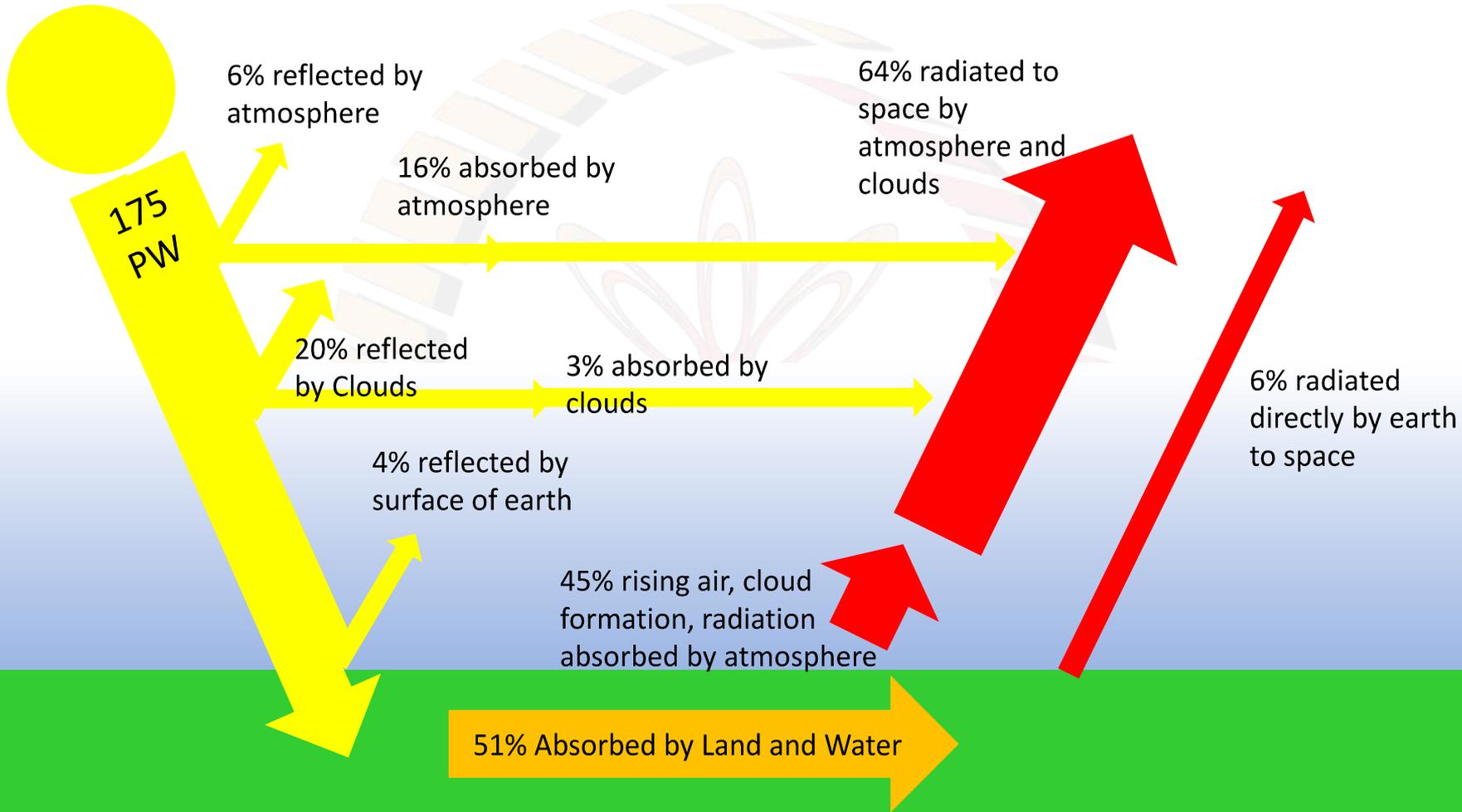


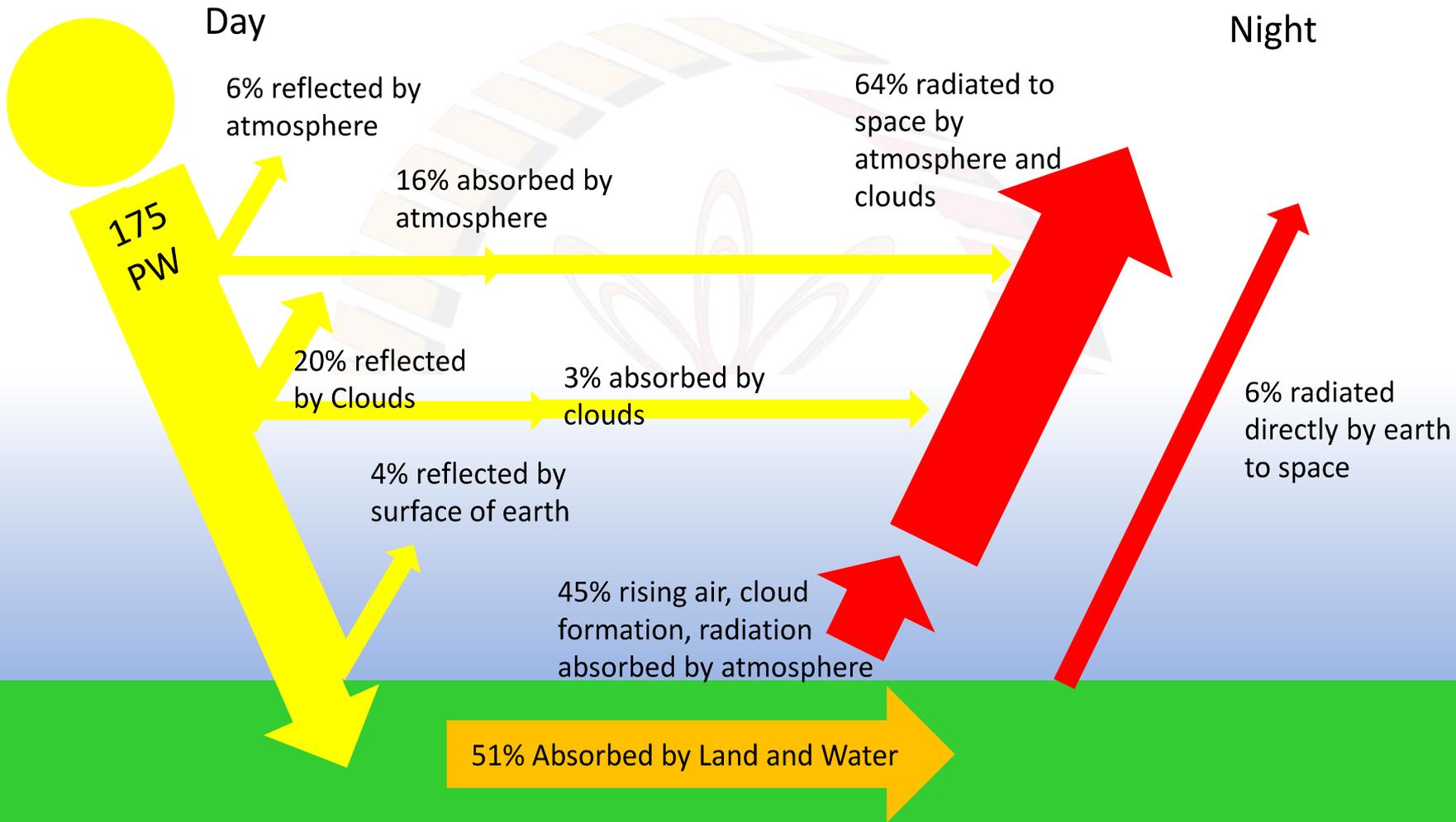


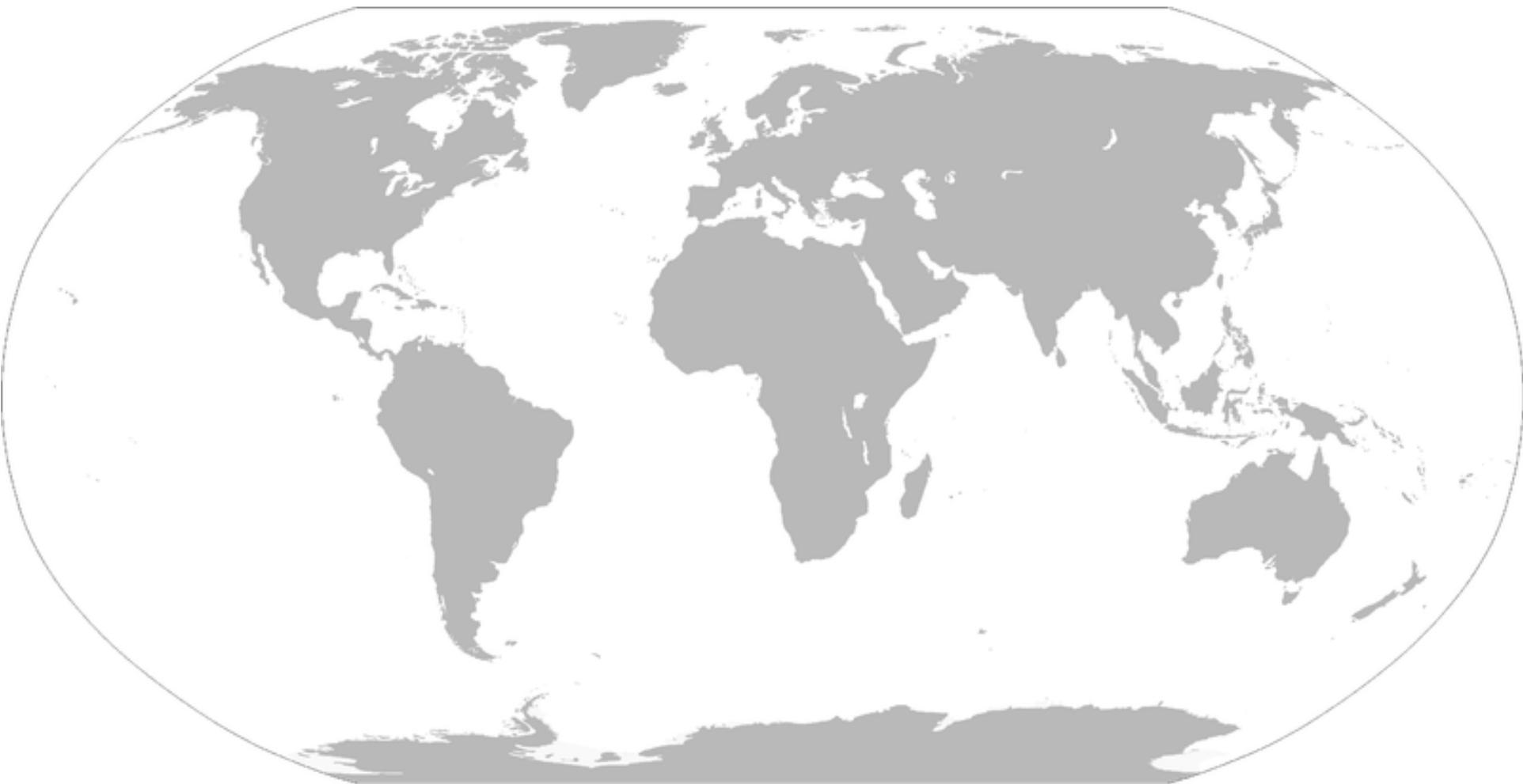






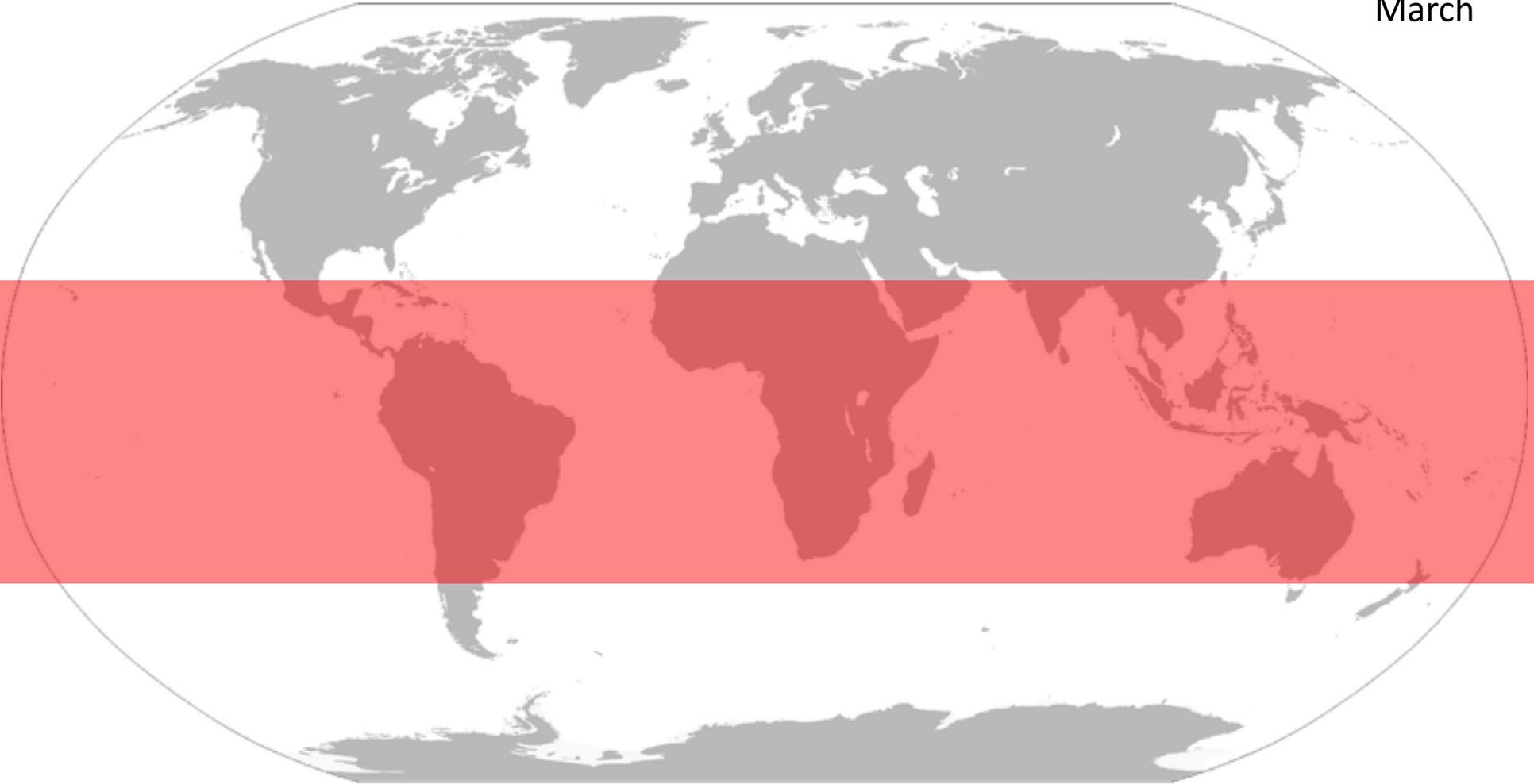






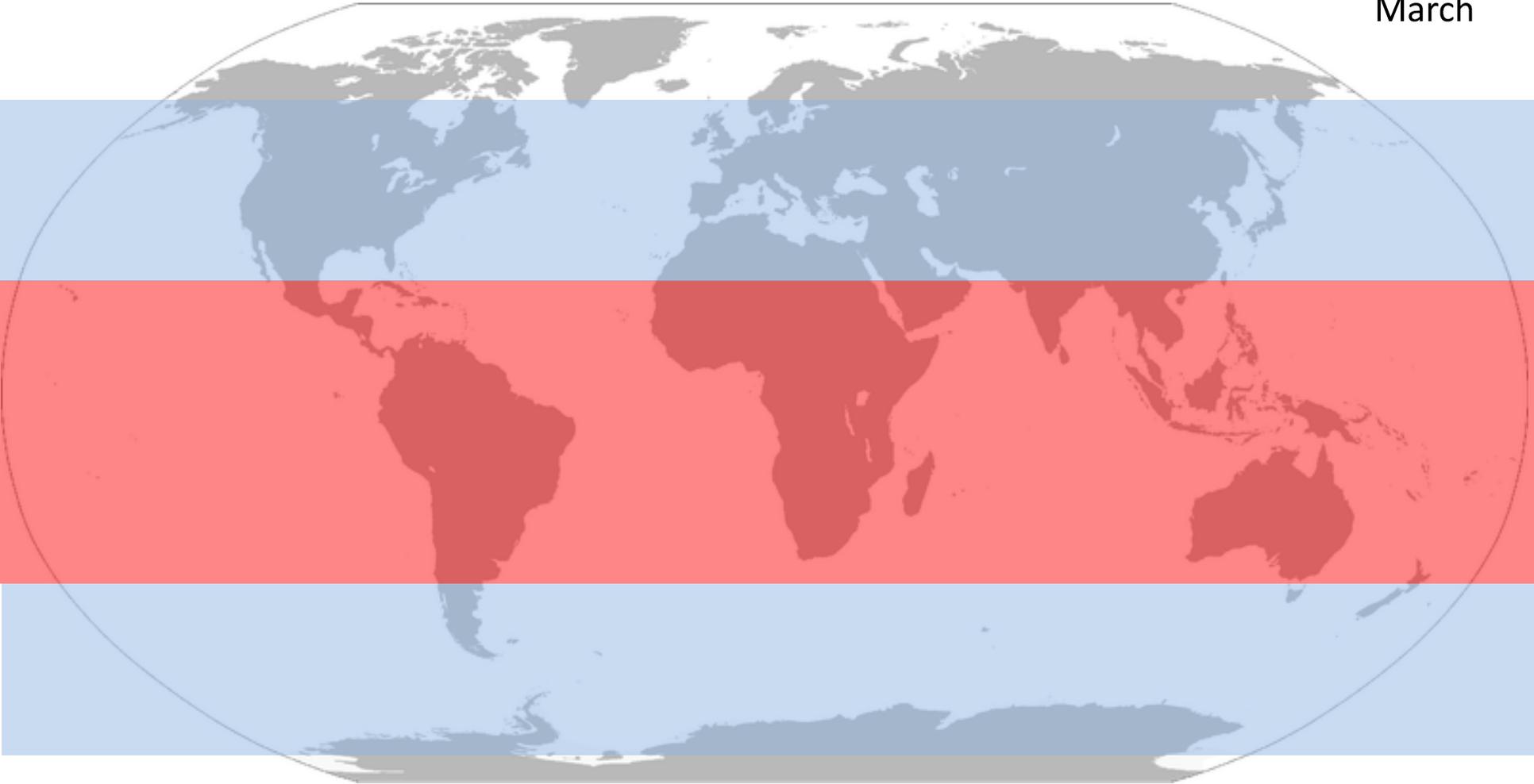
<https://en.wikipedia.org/wiki/File:BlankMap-World-large-noborders.png>

March

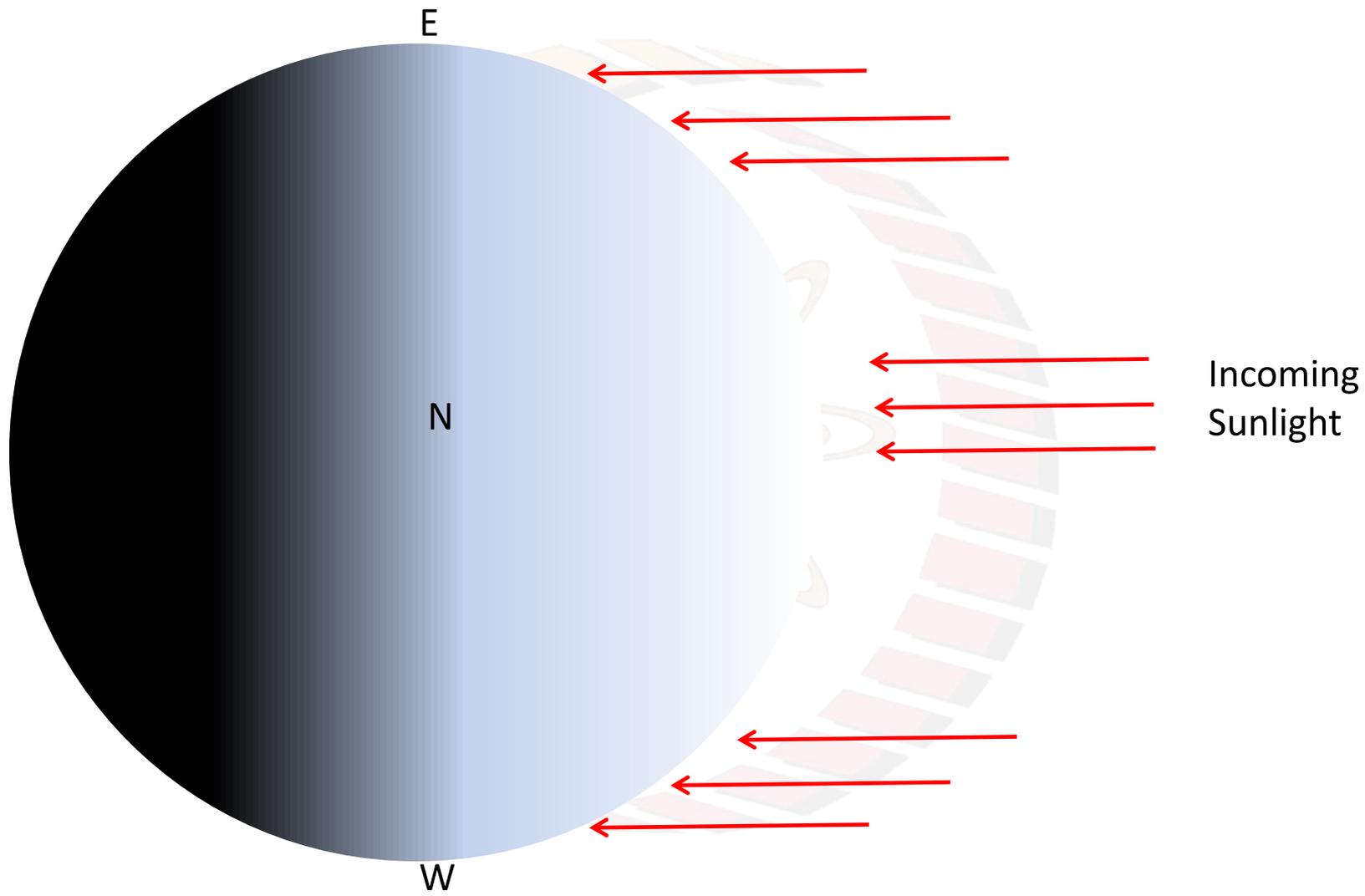


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March



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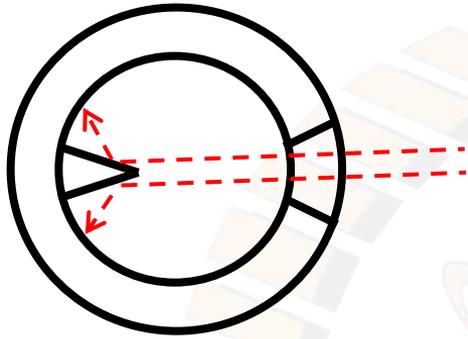


Conclusions:

- 1) Solar energy is absorbed and released through a wide range of phenomena on earth
- 2) Geographical location and seasons are important aspects impacting solar energy received by specific locations
- 3) Time of the day is an important parameter impacting the intensity of solar energy received



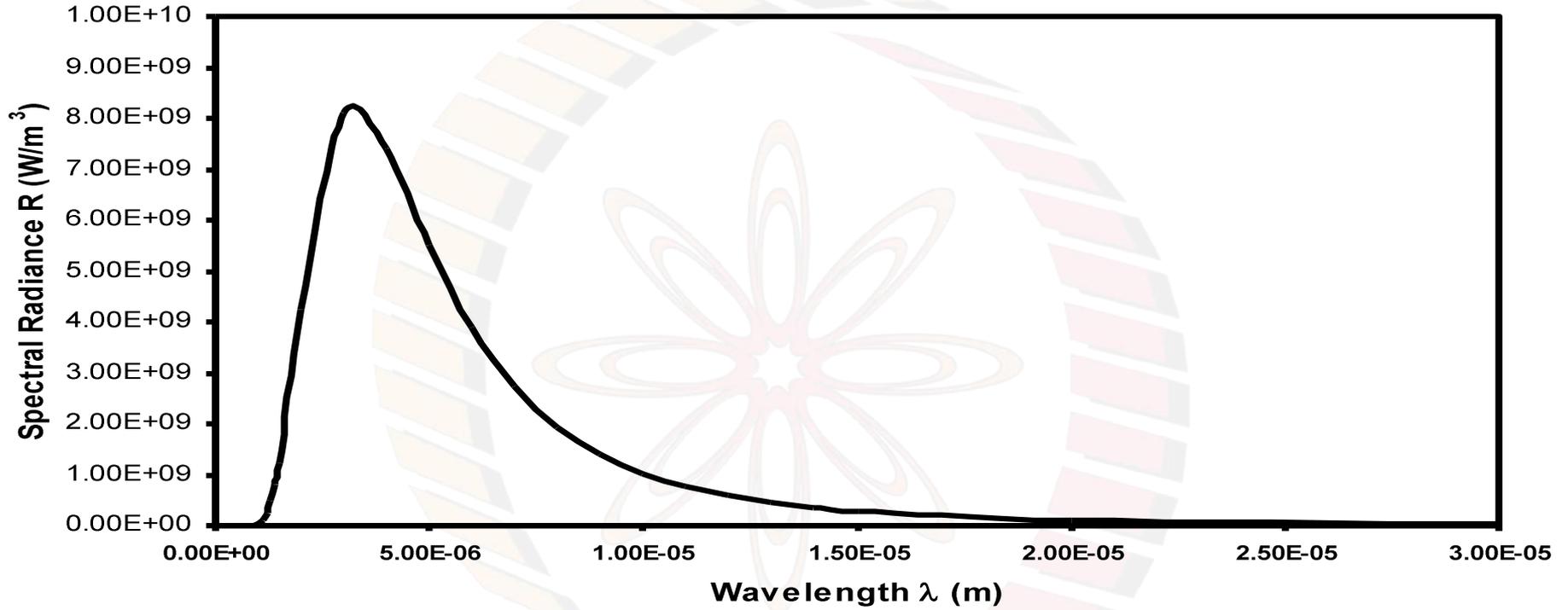
The Solar Spectrum



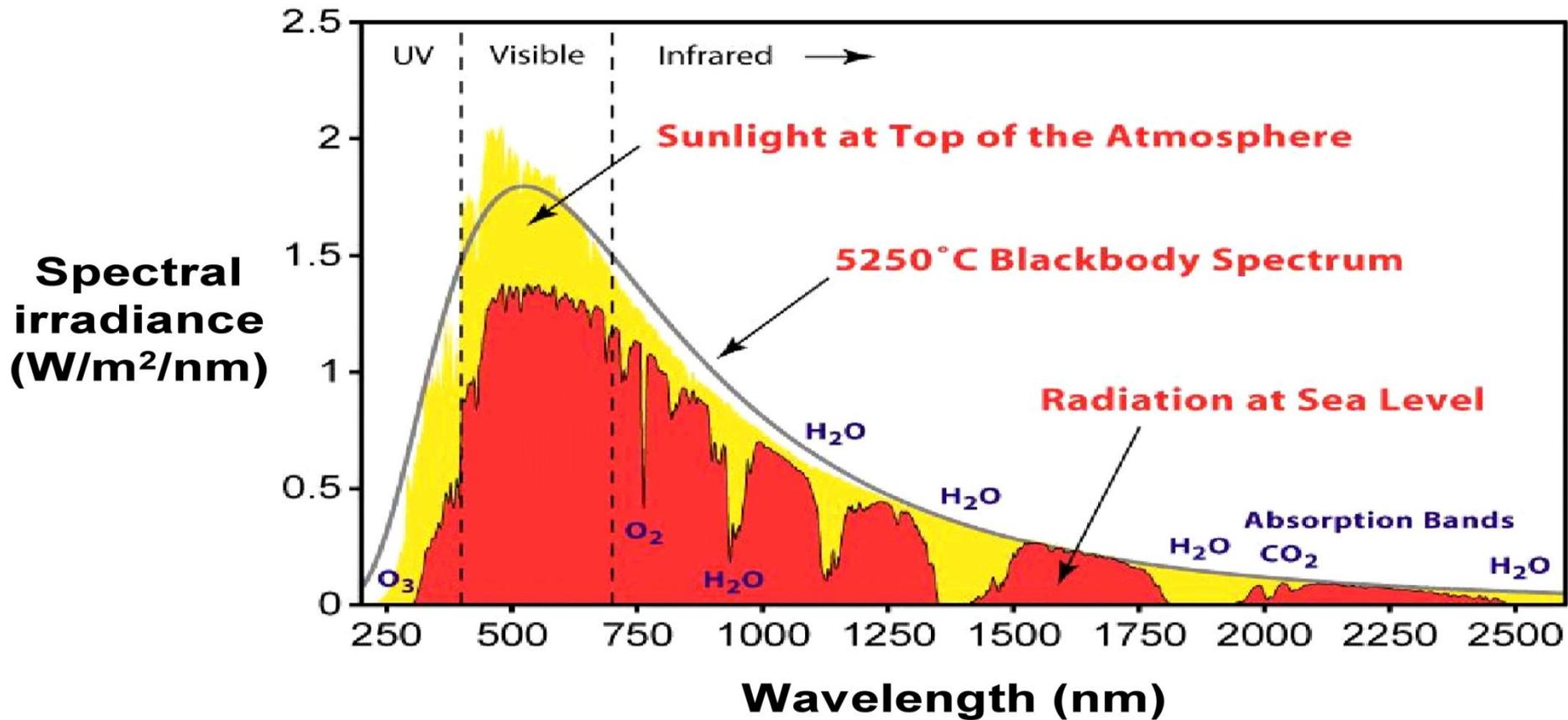
Kirchoff designed a black body in 1859

Known properties of black body radiation:

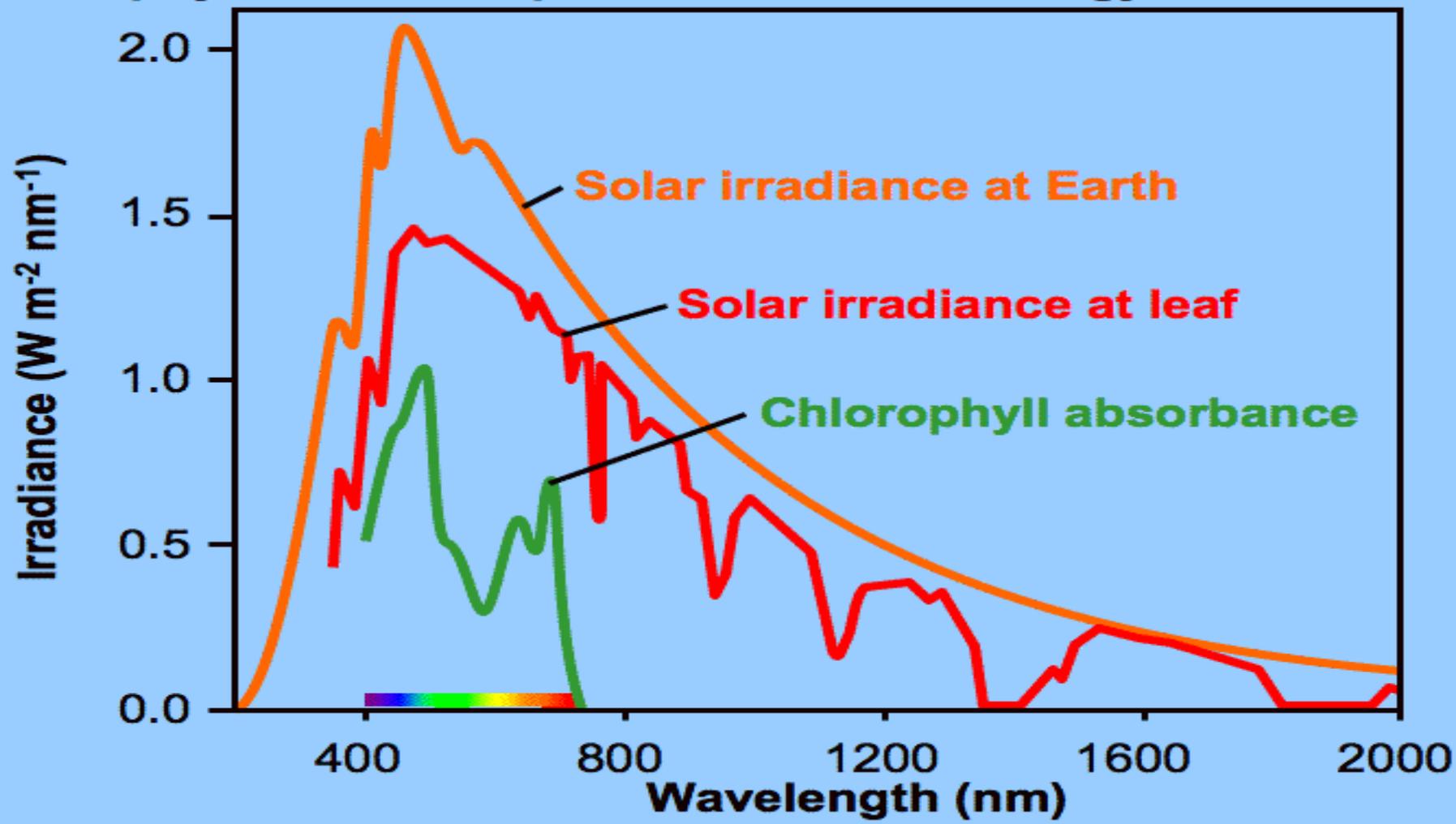
- 1) As temperature T of the body increases, intensity of the radiation from the body also increases**
- 2) Higher the temperature, lower is the wavelength of the most intense part of the spectrum.**

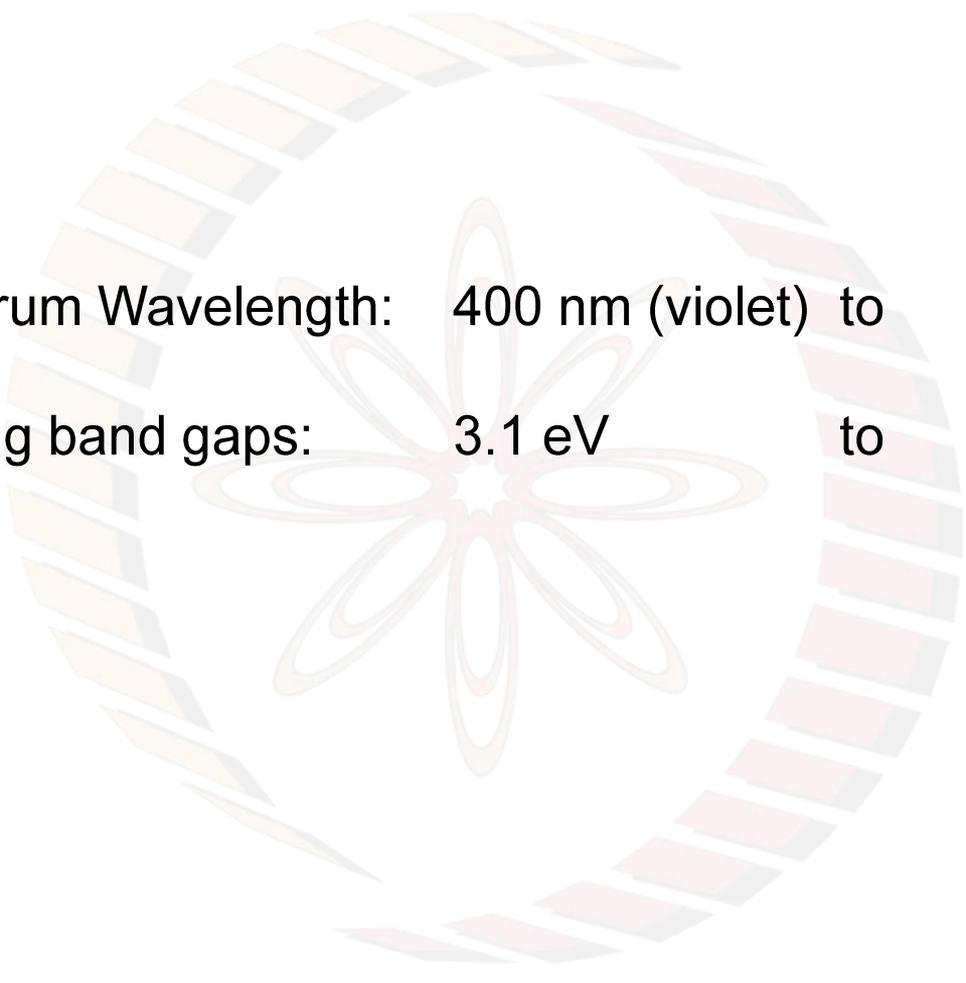


Black body Radiation



Chlorophyll is well-adapted to use Solar Energy



A decorative circular graphic is centered on the page. It features a central floral motif with multiple petals in shades of pink and red. Surrounding this is a ring composed of numerous small, overlapping rectangular segments. The segments on the left side are yellow and orange, while those on the right side are pink and red, creating a color gradient around the circle.

Visible Spectrum Wavelength: 400 nm (violet) to 700 nm (red)
Corresponding band gaps: 3.1 eV to 1.8 eV