

# Solar Energy: The Sun to Earth Transaction

## **Learning objectives:**

- 1) To calculate the energy received by the Earth from the Sun
- 2) To compare the energy received by Earth from the Sun, with the energy usage by humankind

Surface of the Sun  $\sim 5500\text{ }^{\circ}\text{C}$

Core of the Sun, several million  $^{\circ}\text{C}$

**Sun gives out 384 Yotta Watts**

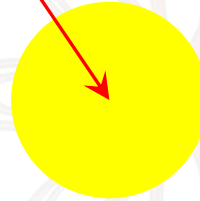
$$= 384 \times 10^{24} \text{ W}$$

$$= 3.84 \times 10^{26} \text{ W}$$

Earth's radius:  
6371 km

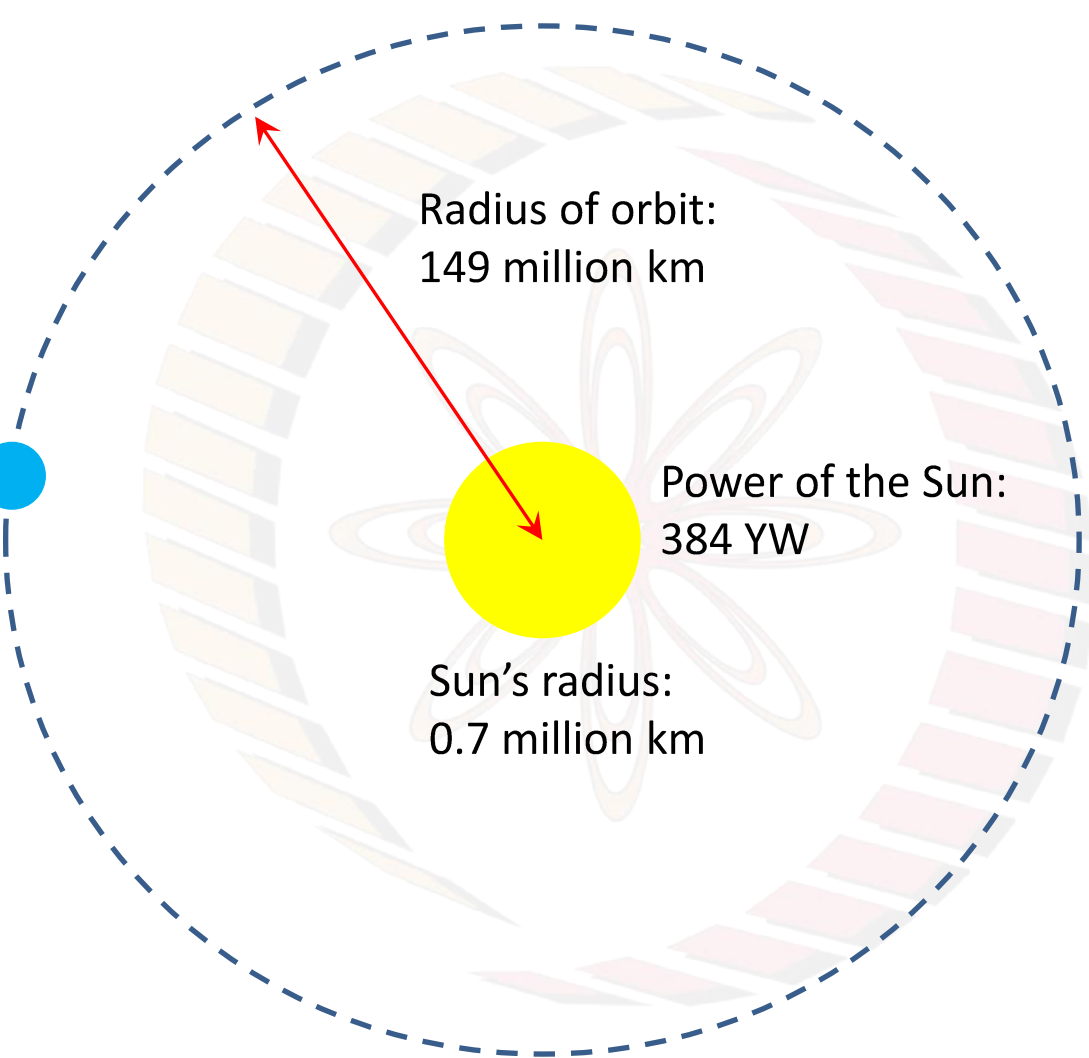


Radius of orbit:  
149 million km



Sun's radius:  
0.7 million km

Power of the Sun:  
384 YW

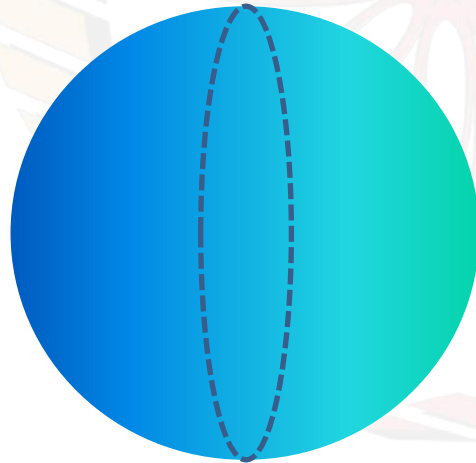


## Intensity of Sun's radiation at Earth's orbit:

$$= \frac{3.84 \times 10^{26}}{4 \times 3.14 \times (1.49 \times 10^{11})^2} = 1377 \text{ W/m}^2$$

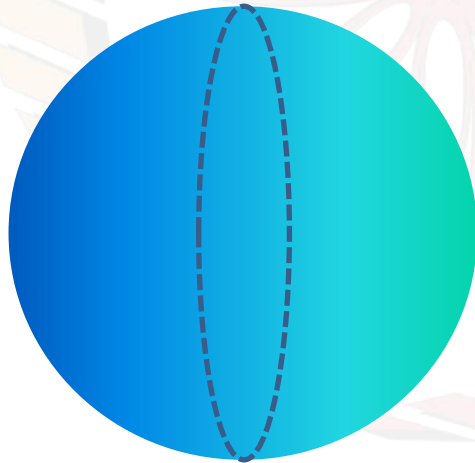
**Area of Earth's disc:**

$$= 3.14 \times (6.371 \times 10^6)^2 = 1.27 \times 10^{14} \text{ m}^2$$



## Power received from the Sun, by Earth:

$$= 1.27 \times 10^{14} \times 1377 = 1.755 \times 10^{17} \text{ W or J/s}$$



**Energy received from the Sun, by Earth each year:**

$$= 1.755 \times 10^{17} \times 60 \times 60 \times 24 \times 365$$

$$= 5.5 \times 10^{24} \text{ J}$$

**= 5.5 million Exa Joules per year**





**Humankind uses:**

= 500 Exa Joules per year

**Earth receives from the Sun:**

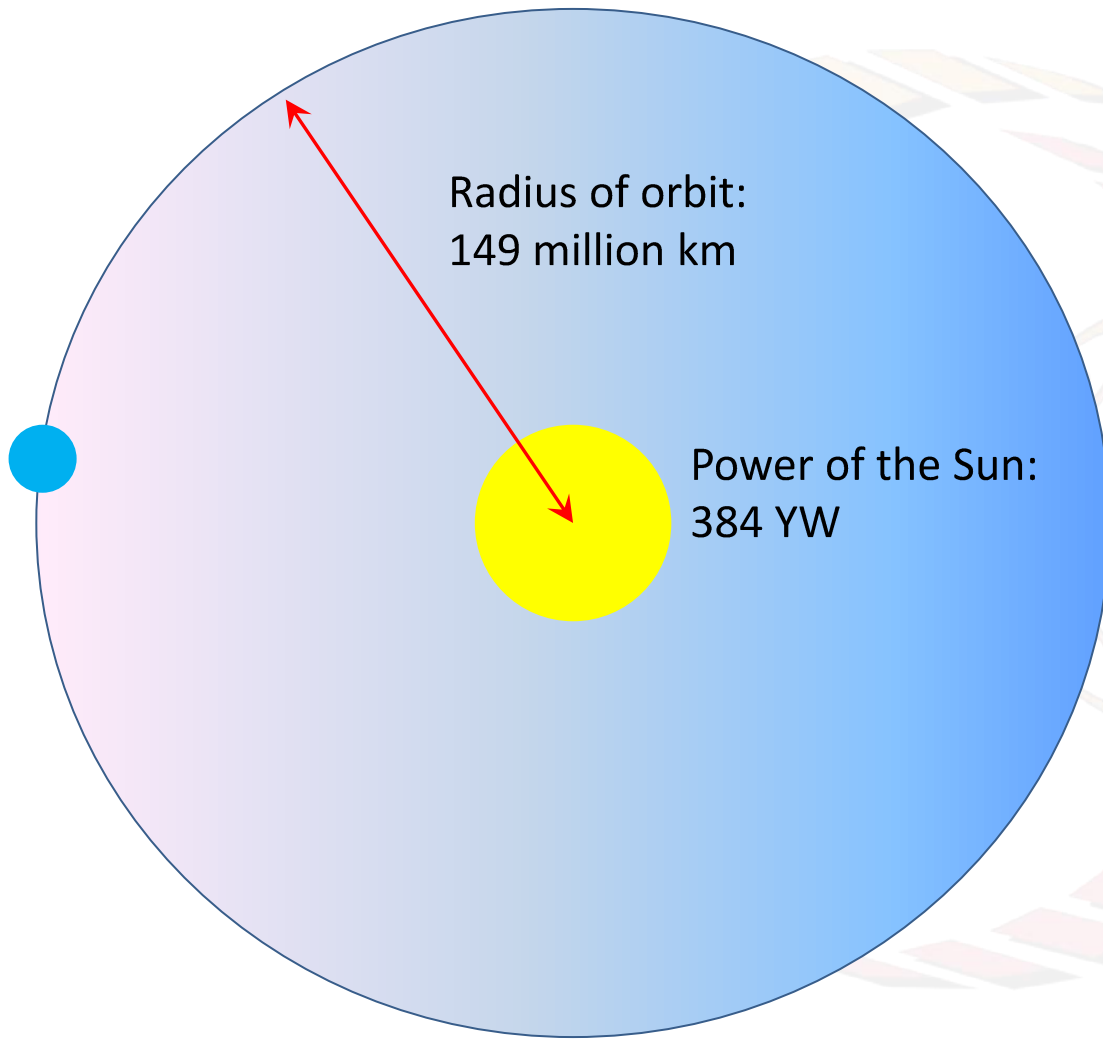
5.5 million Exa Joules per year

**This is received in :**

$500 / 5.5 \times 10^6 = 9 \times 10^{-5} \text{ years} = 0.033 \text{ days} = 0.79 \text{ hours}$

Since 30% of the incident energy is reflected back, on the surface of the Earth, the energy used by humankind each year is received in :

$$= 0.79/0.70 \sim \mathbf{1 \text{ hour}}$$



## **Dyson Sphere:**

Freeman Dyson (1960)

## **Kardashev Scale (1964)**


Type 1: All energy reaching planet

Type 2: All energy of Star

Type 3: All energy of Galaxy

## Conclusions:

- 1) Earth receives nearly 5.5 million exajoules of energy from the Sun each year
- 2) The entire energy used by humankind each year, is received on the surface of the earth, from the sun, each hour!

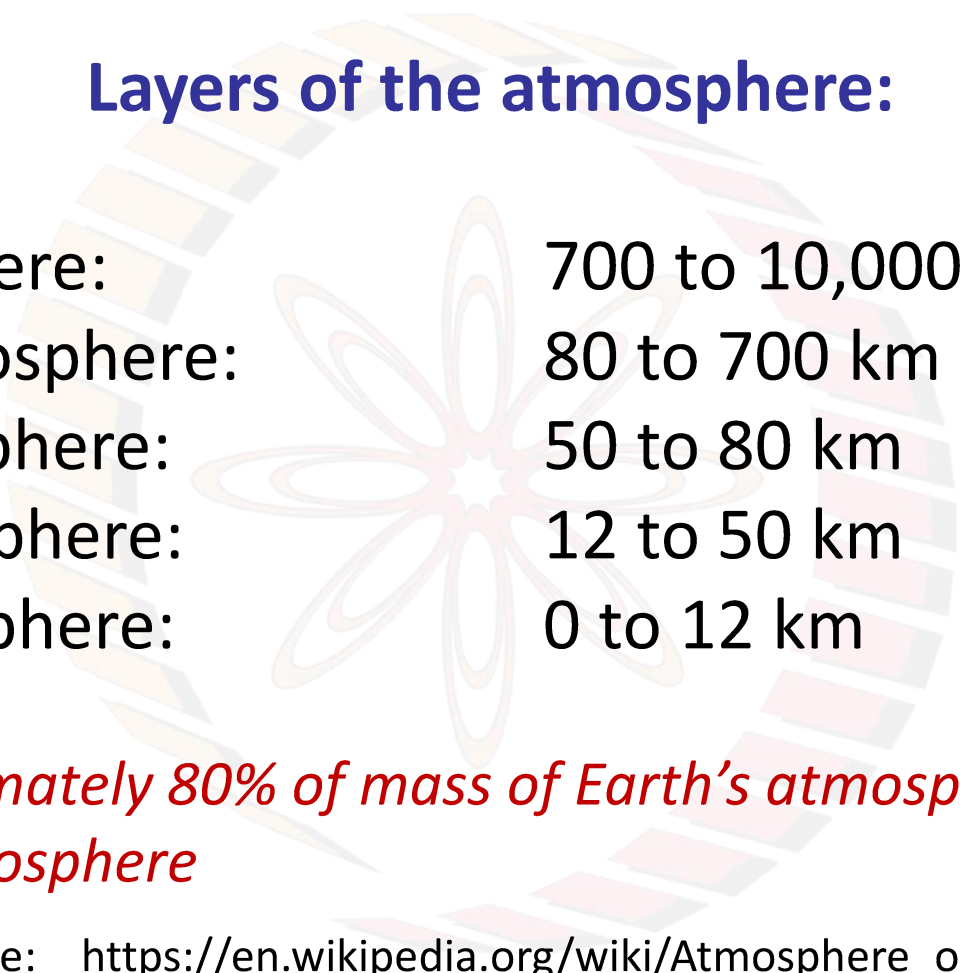


# The Solar Energy Budget

## Learning objectives:

- 1) To know how incoming solar energy is distributed across various phenomena on earth
- 2) To become aware of geographical and seasonal aspects associated with solar energy
- 3) To become aware of impact of time of day on ability to receive solar energy

## Layers of the atmosphere:



|               |                  |
|---------------|------------------|
| Exosphere:    | 700 to 10,000 km |
| Thermosphere: | 80 to 700 km     |
| Mesosphere:   | 50 to 80 km      |
| Stratosphere: | 12 to 50 km      |
| Troposphere:  | 0 to 12 km       |

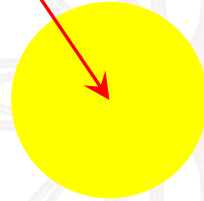
*Approximately 80% of mass of Earth's atmosphere is in the Troposphere*

Source: [https://en.wikipedia.org/wiki/Atmosphere\\_of\\_Earth](https://en.wikipedia.org/wiki/Atmosphere_of_Earth)

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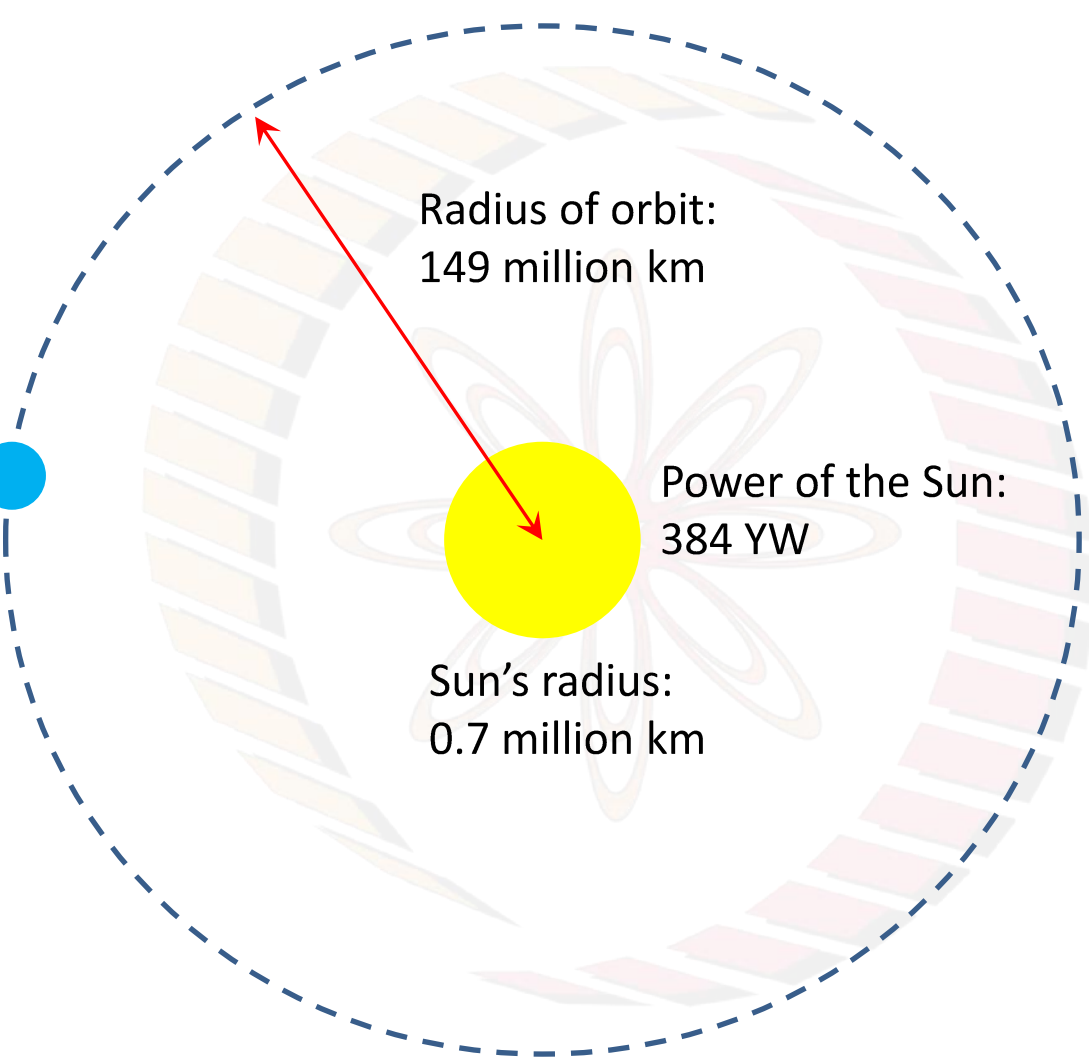


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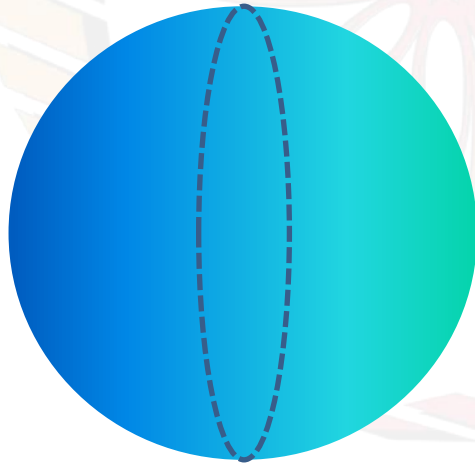
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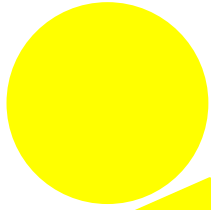




## Power received from the Sun, by Earth:

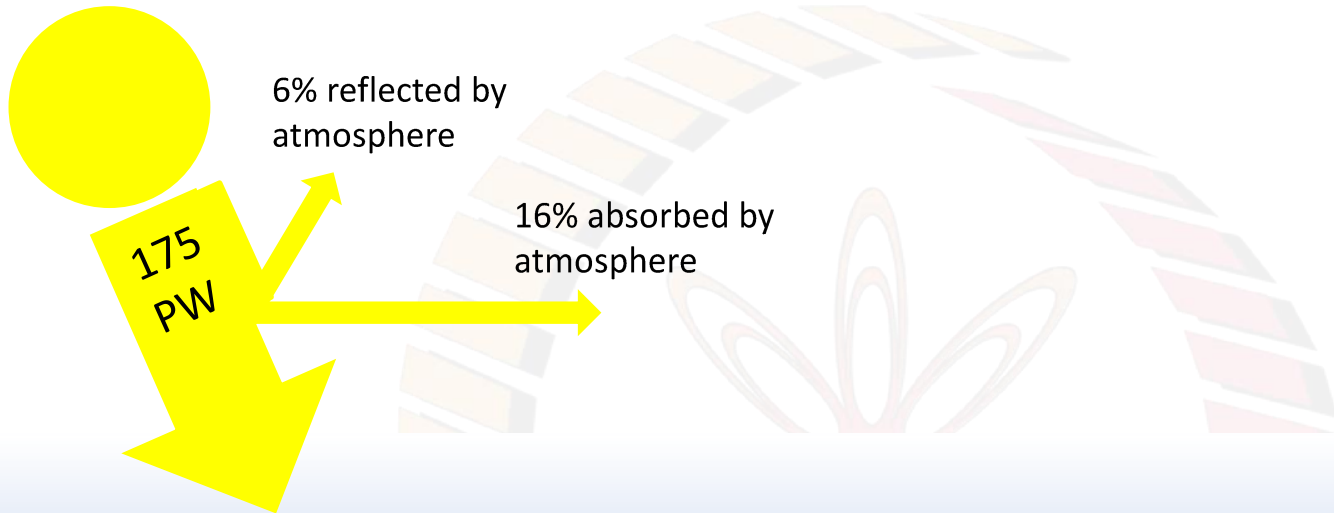
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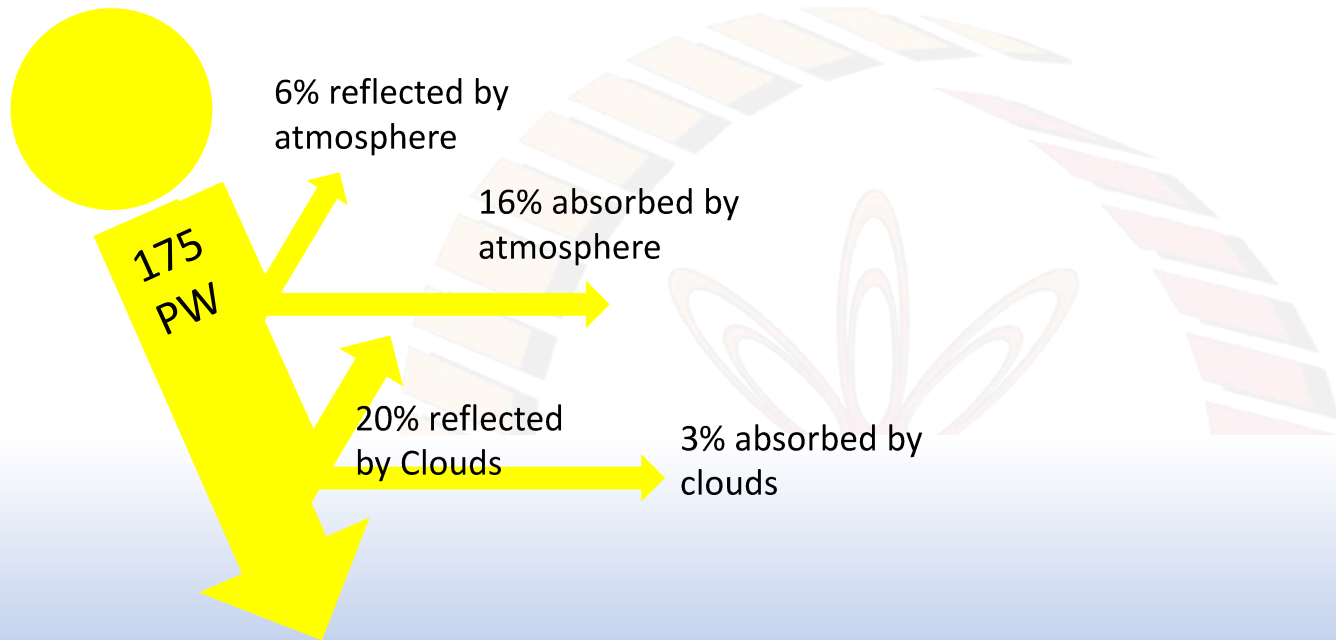


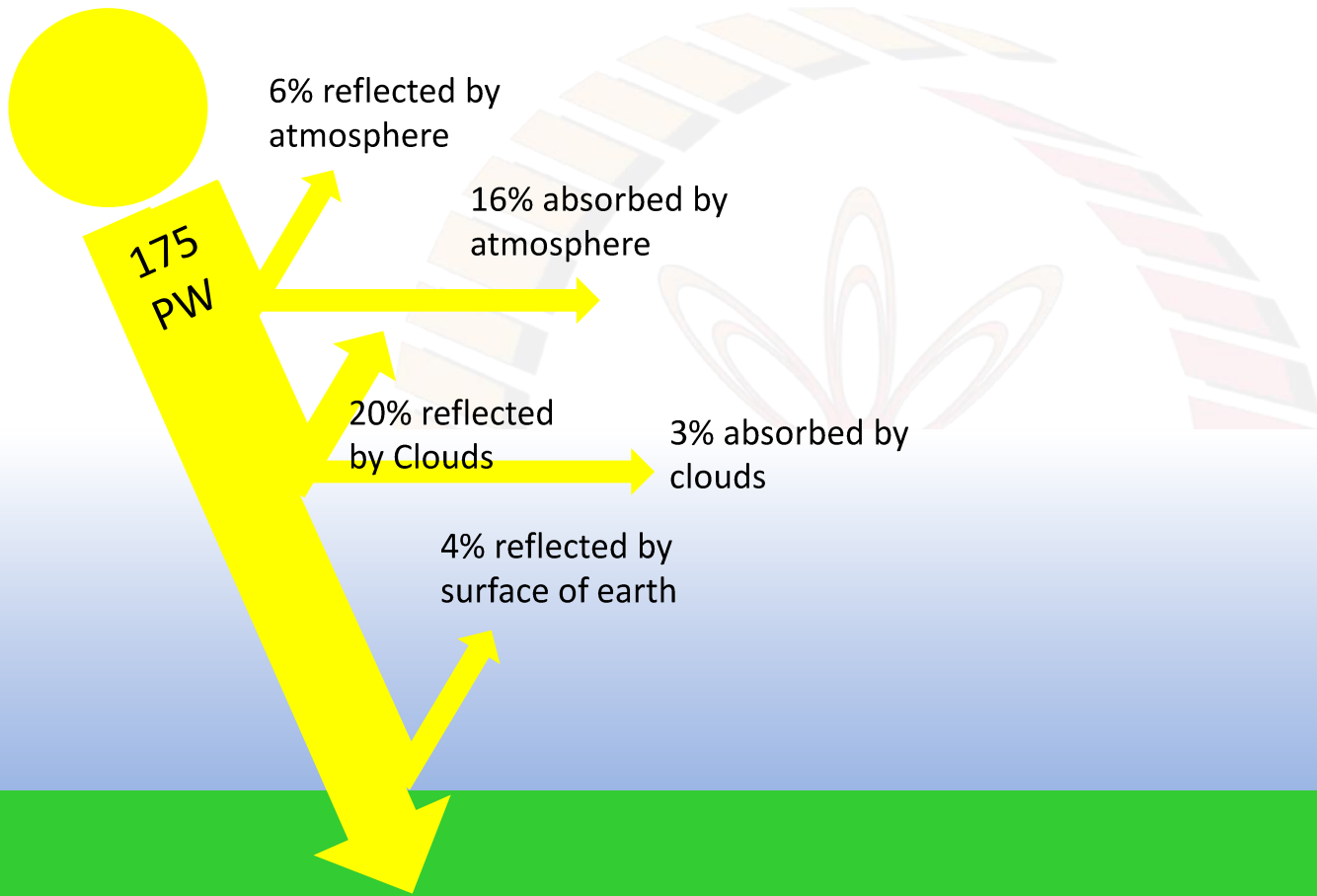


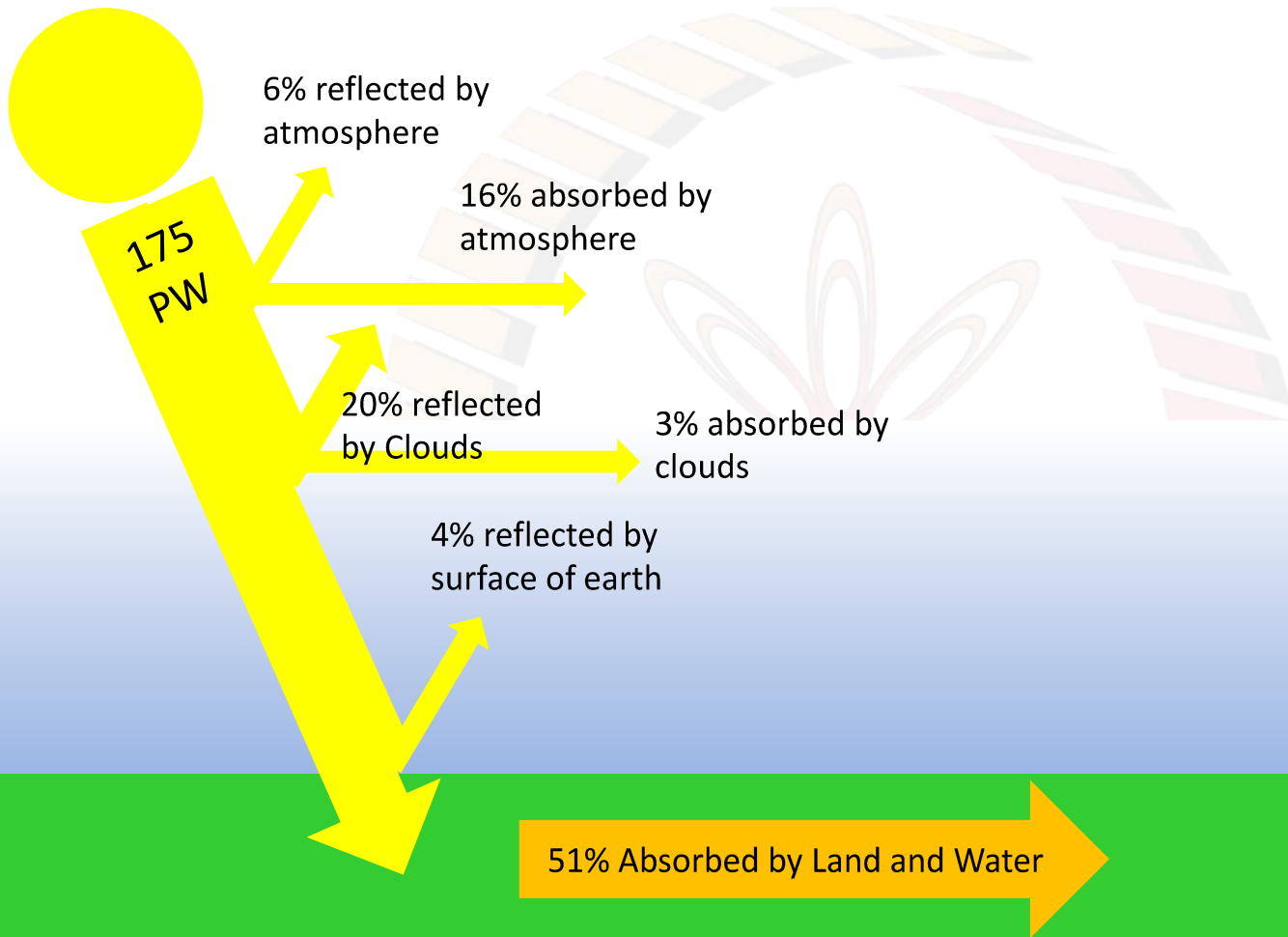
175  
PW

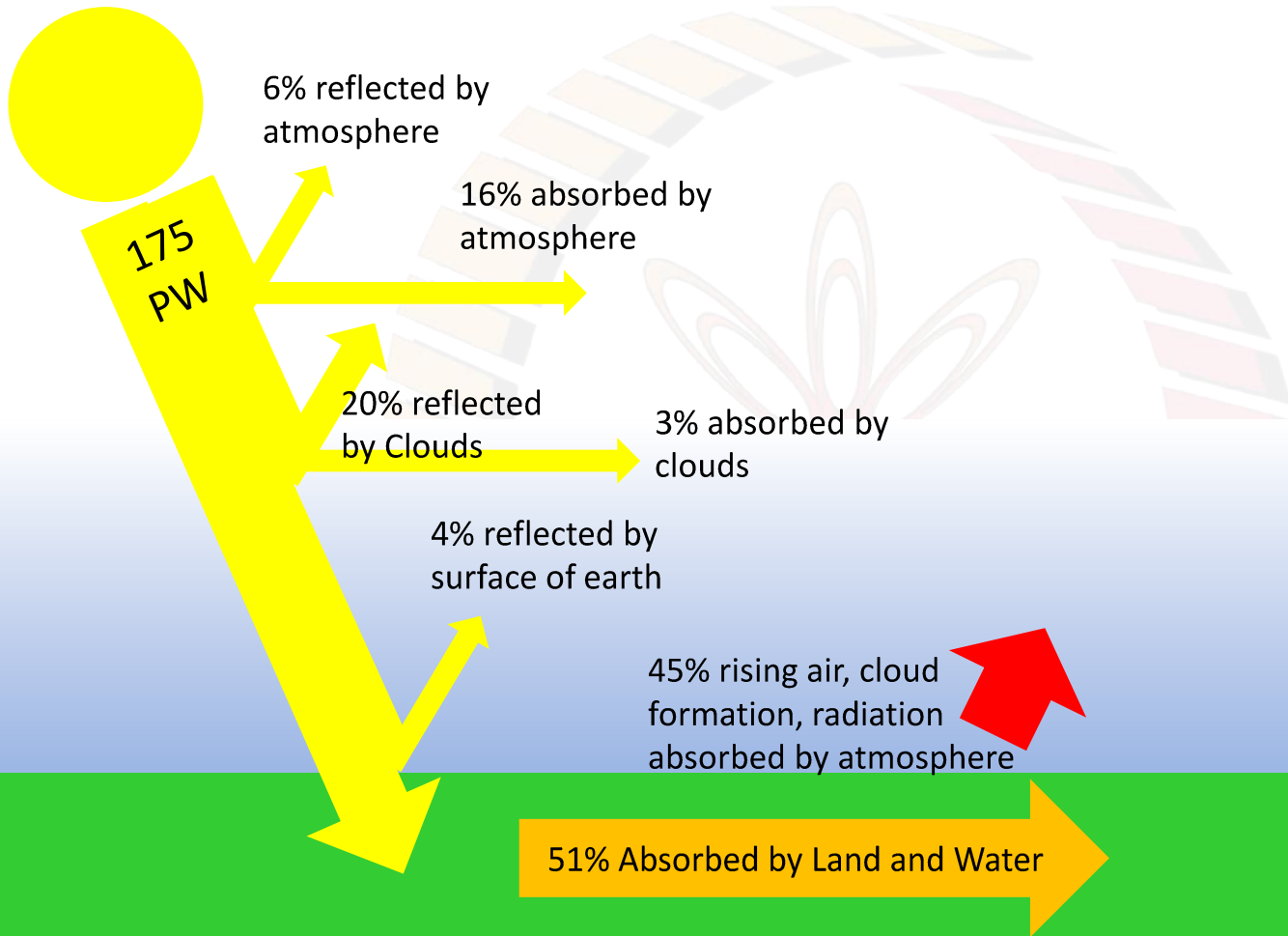


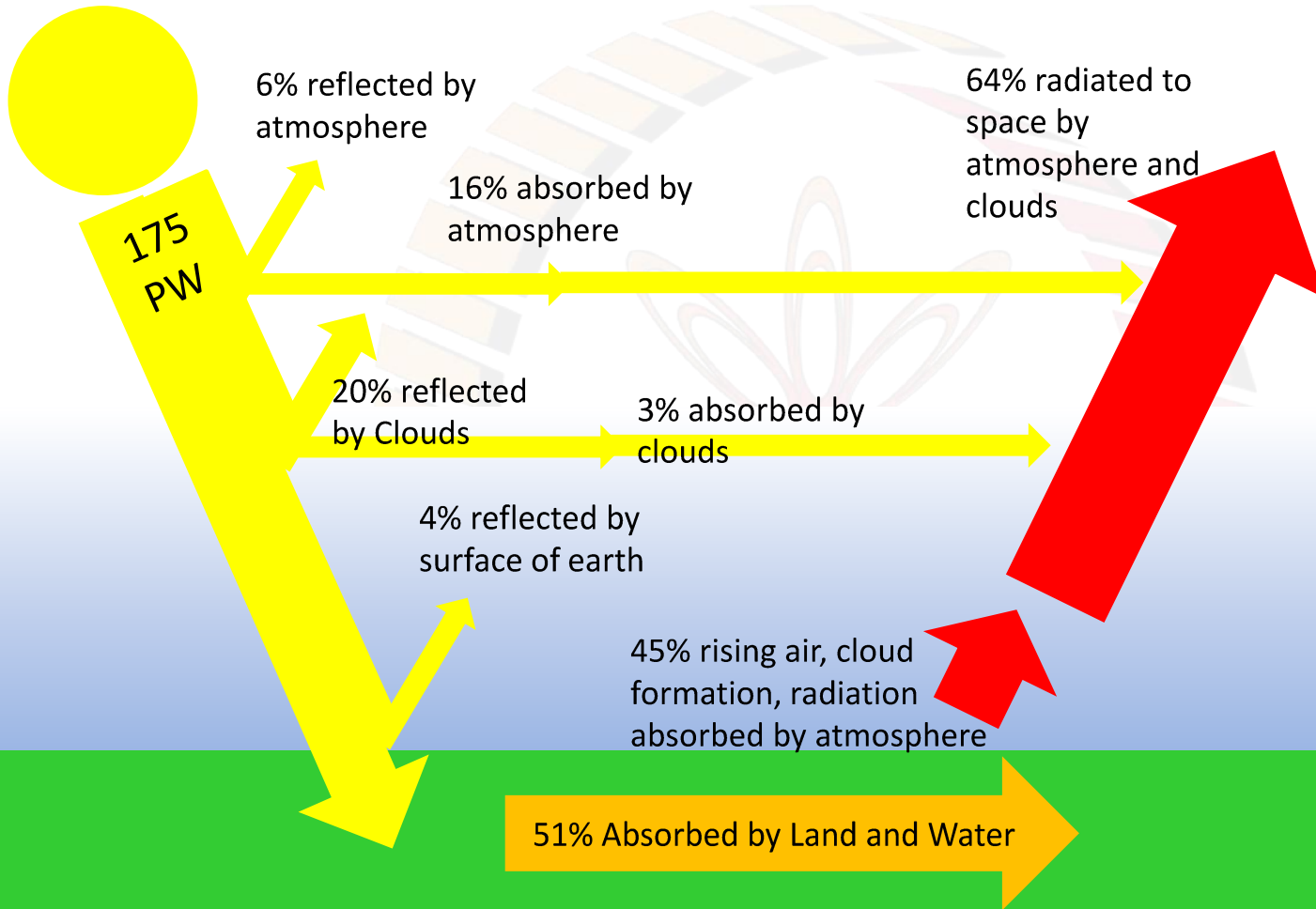




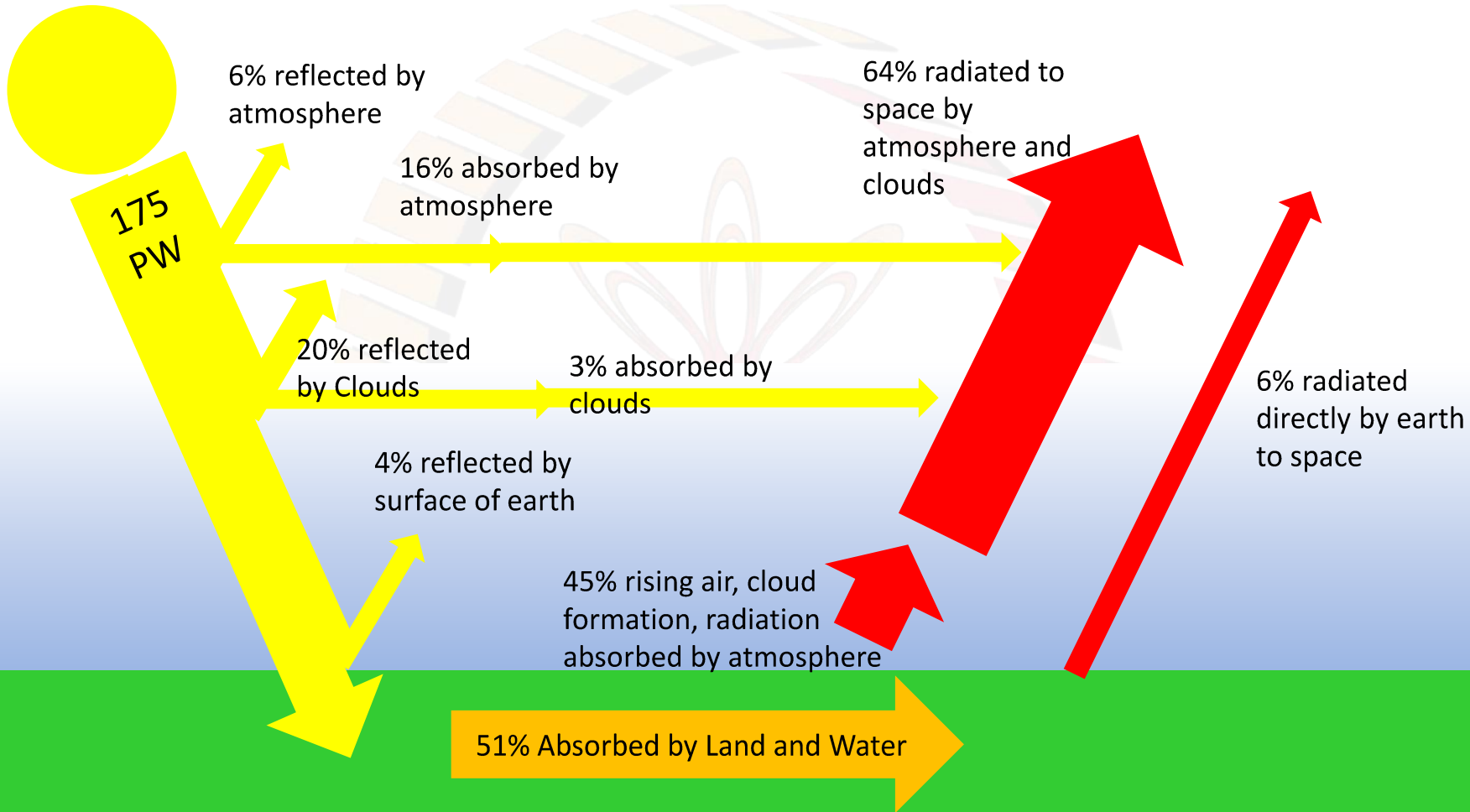


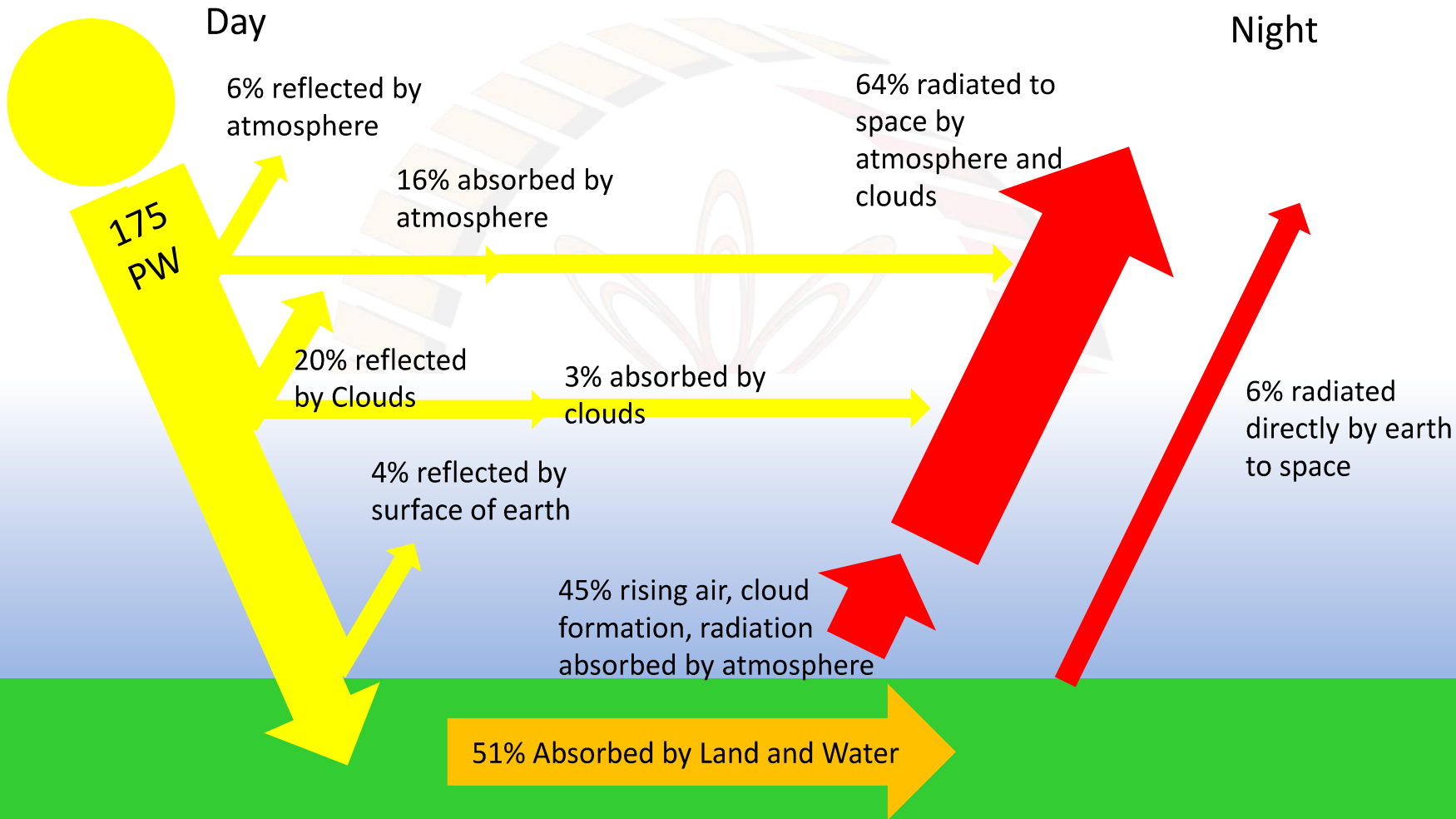


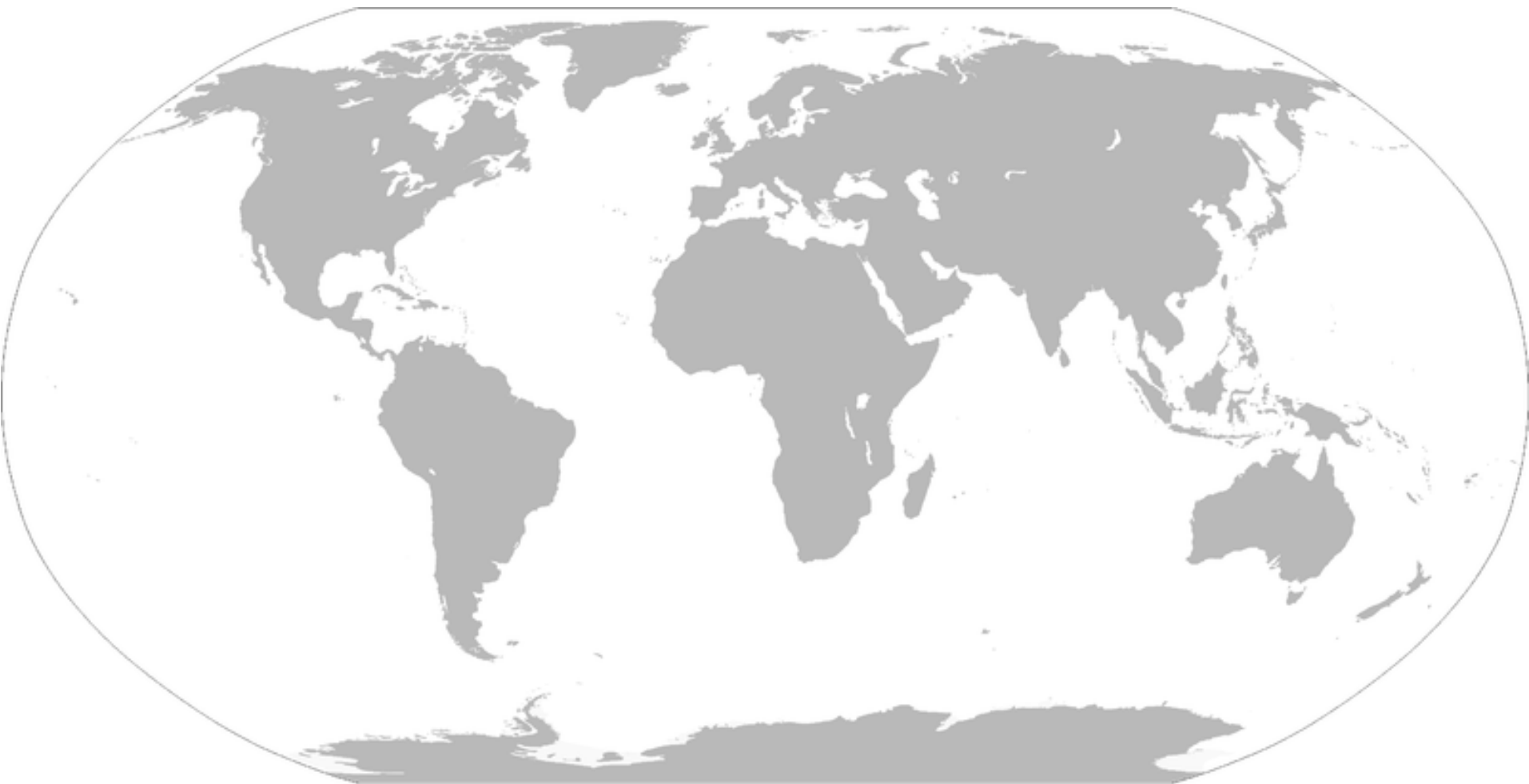






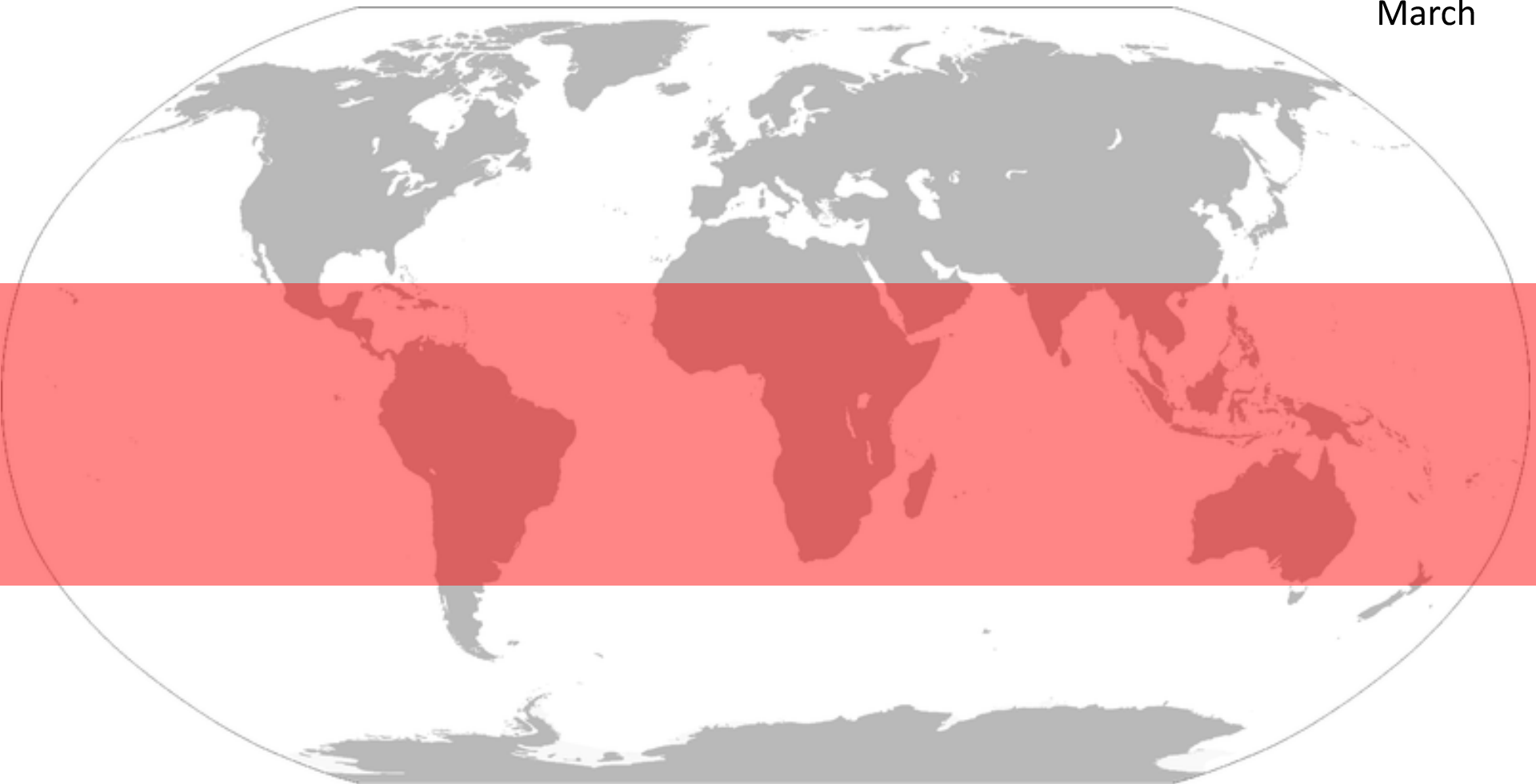






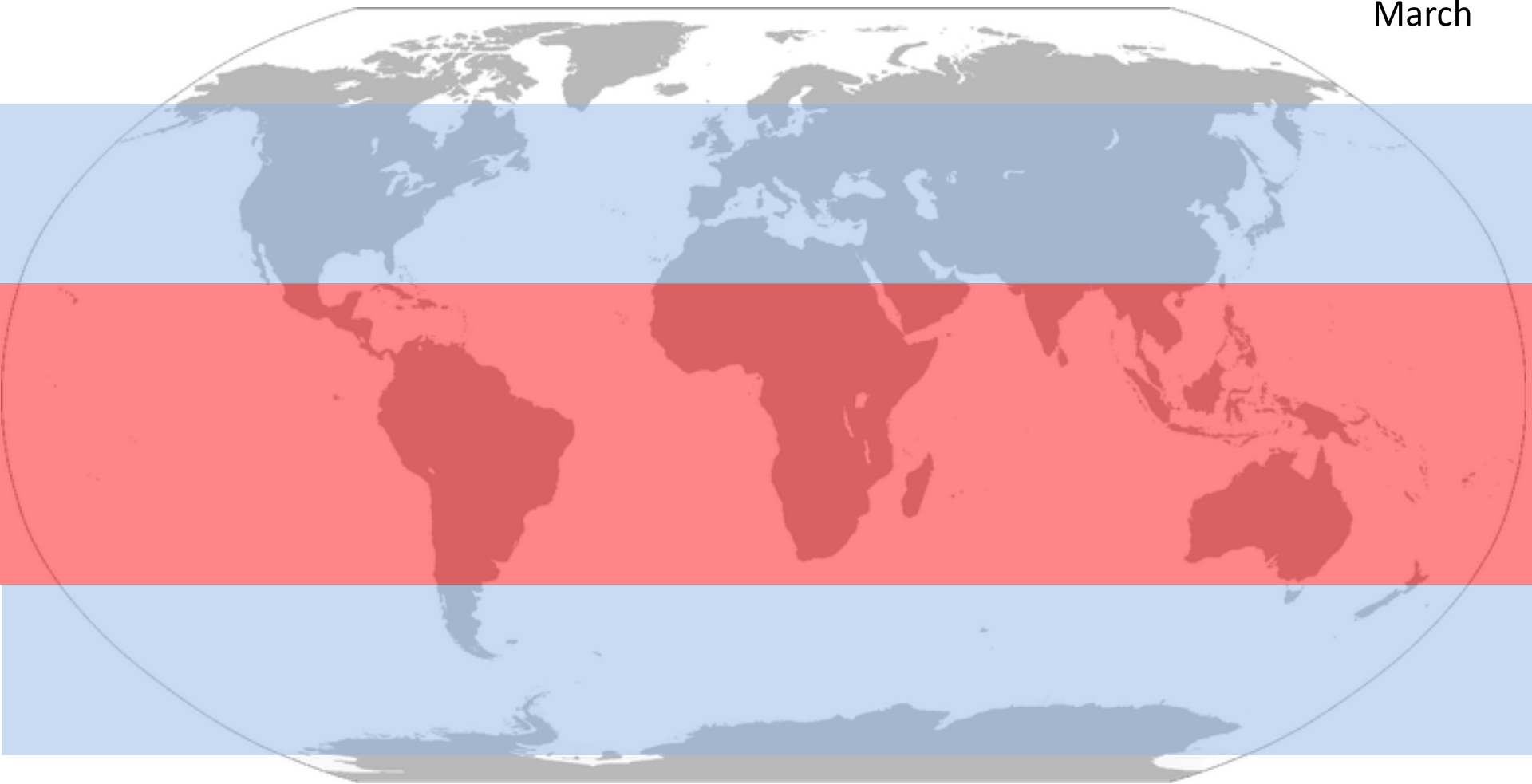
<https://en.wikipedia.org/wiki/File:BlankMap-World-large-noborders.png>

March

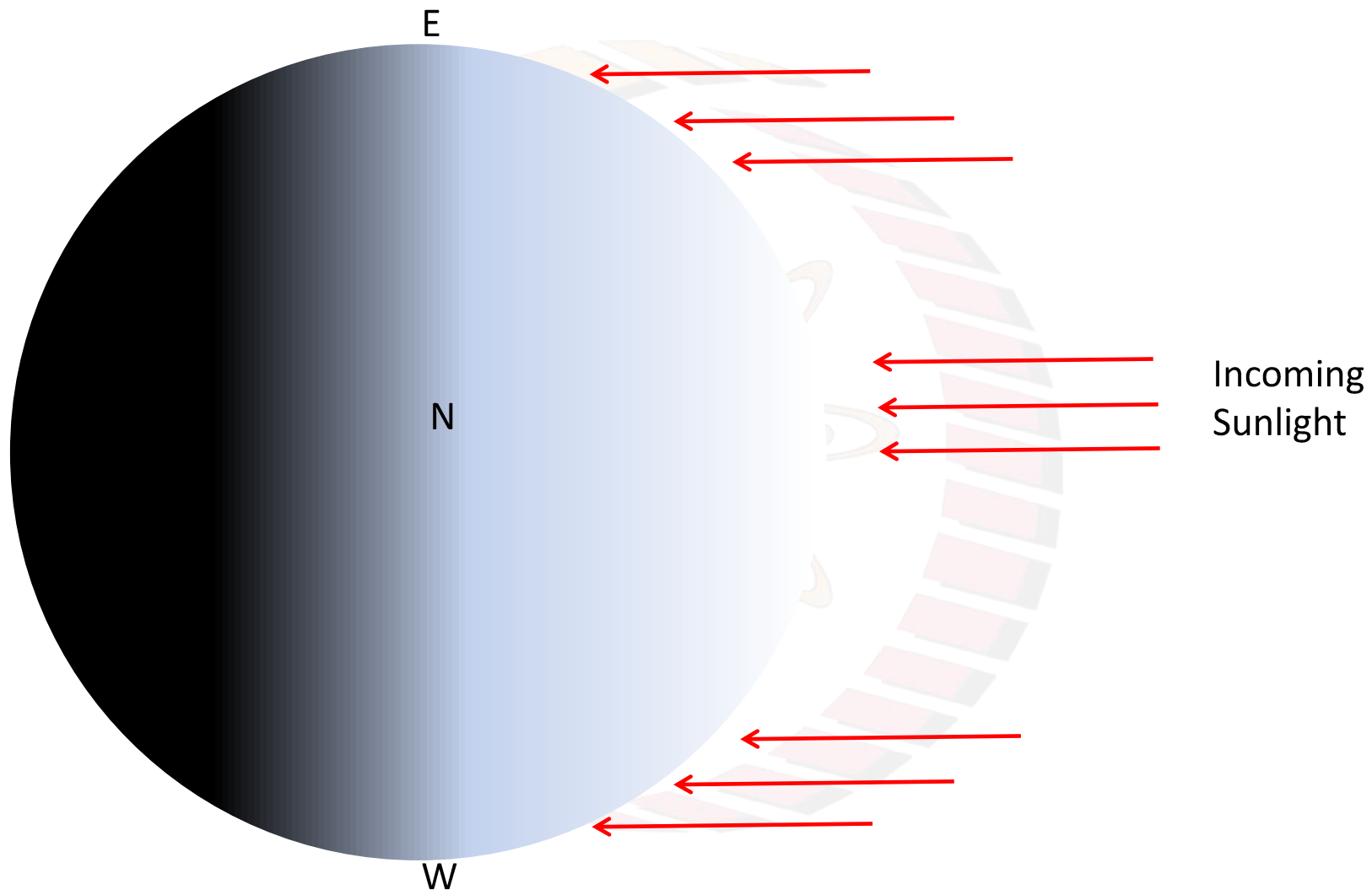


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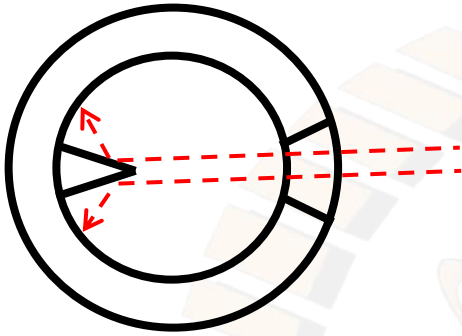
## Conclusions:

- 1) Solar energy is absorbed and released through a wide range of phenomena on earth
- 2) Geographical location and seasons are important aspects impacting solar energy received by specific locations
- 3) Time of the day is an important parameter impacting the intensity of solar energy received

# The Solar Spectrum

The background of the slide is a light gray. In the center, there is a large, faint, circular graphic. This graphic consists of a ring of rectangular blocks arranged in a circle. The blocks are colored in a gradient: yellow on the left, orange at the top, red on the right, and pink at the bottom. In the center of this ring is a stylized flower or star shape with eight petals or points, colored in a gradient from light yellow to light pink.

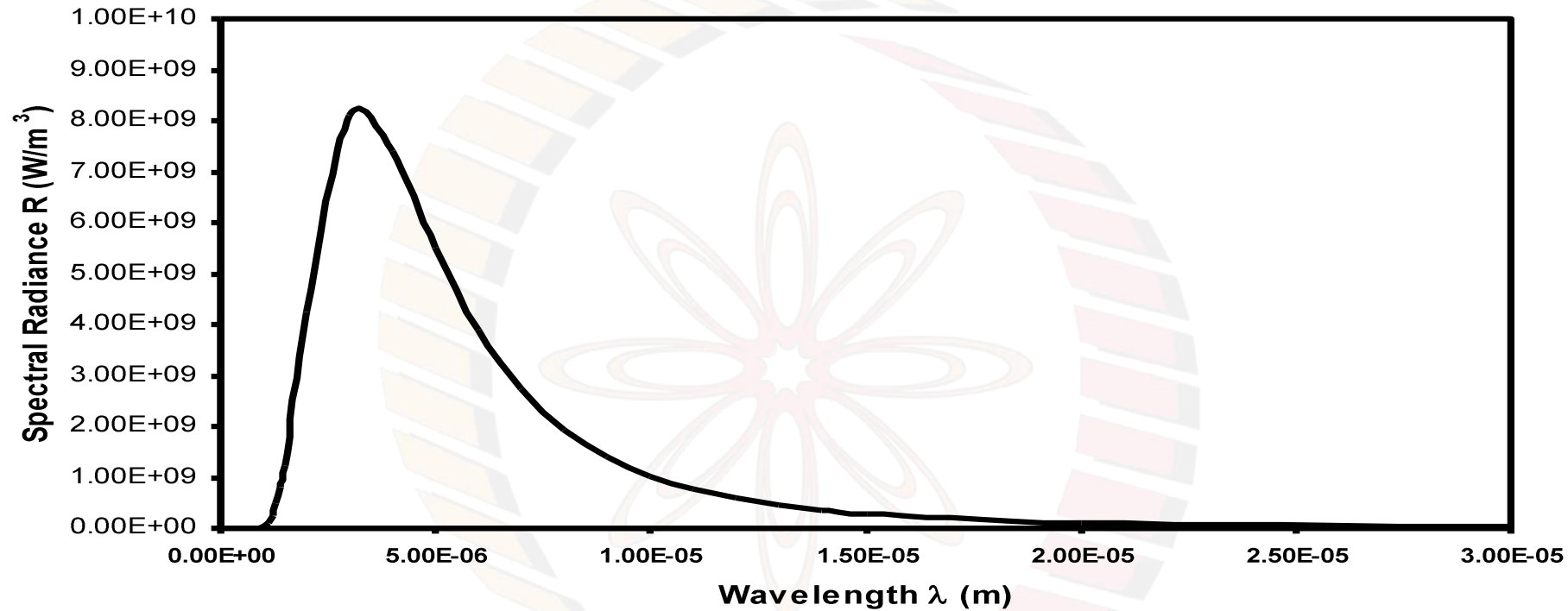




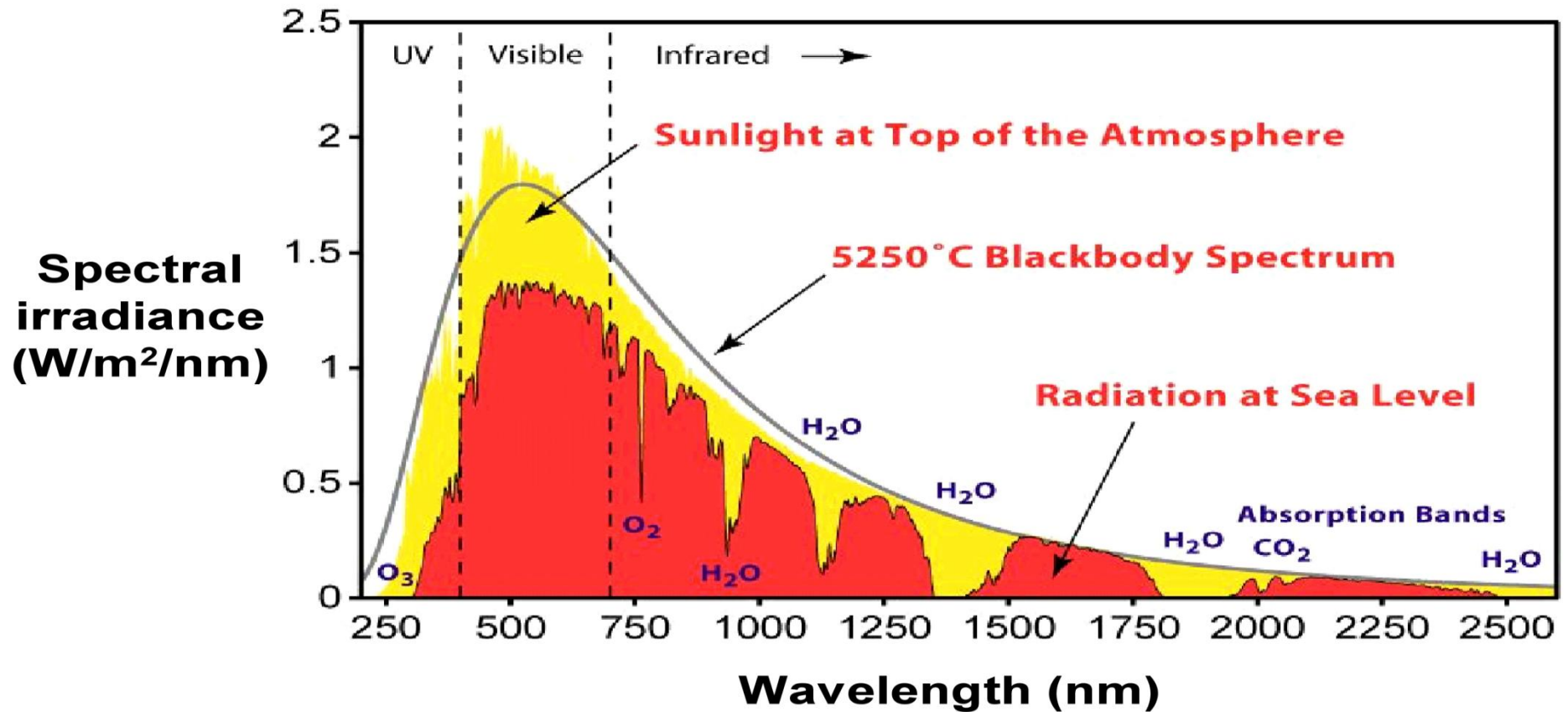
**Kirchoff** designed a black body in 1859

### **Known properties of black body radiation:**

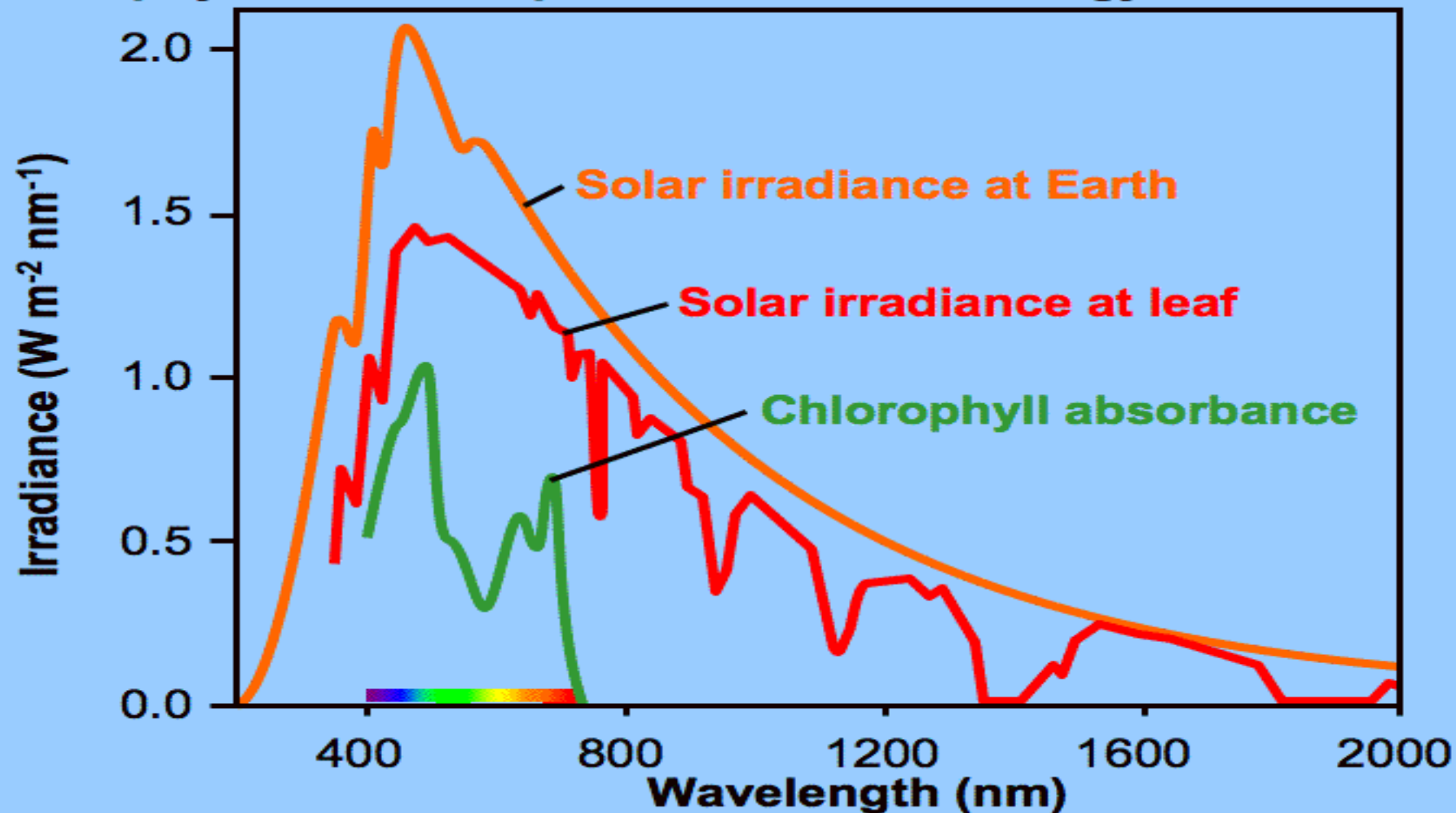
- 1) As temperature  $T$  of the body increases, intensity of the radiation from the body also increases**
- 2) Higher the temperature, lower is the wavelength of the most intense part of the spectrum.**



**Black body Radiation**



## Chlorophyll is well-adapted to use Solar Energy





Visible Spectrum Wavelength: 400 nm (violet) to 700 nm (red)

Corresponding band gaps: 3.1 eV to 1.8 eV